The Spatial Distribution of Upstream Ion Events Measured by ACE, Wind, and STEREO-A Near The Earth's Bow Shock

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Big Issues: Upstream Ion Events

• Where accelerated?

→ Ring current, cusp, magnetotail, bow shock

• What material?

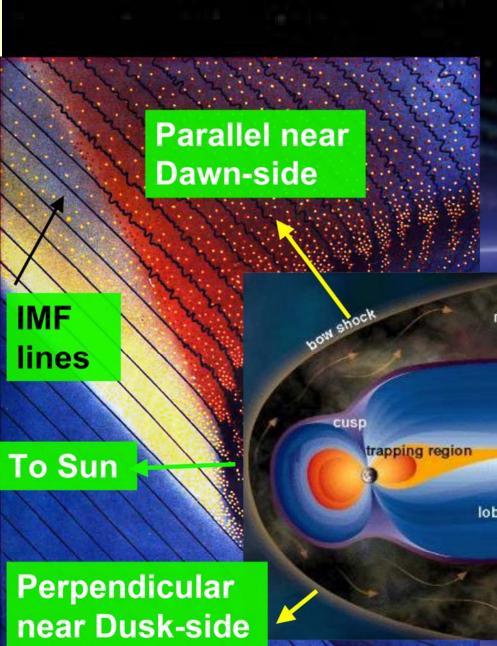
→ Ionospheric ions, solar wind, or other

• How accelerated?

→ Reconnection-driven, ULF-waves, parallel E fields, or bow-shock acceleration

Transport

→ Leakage, streaming

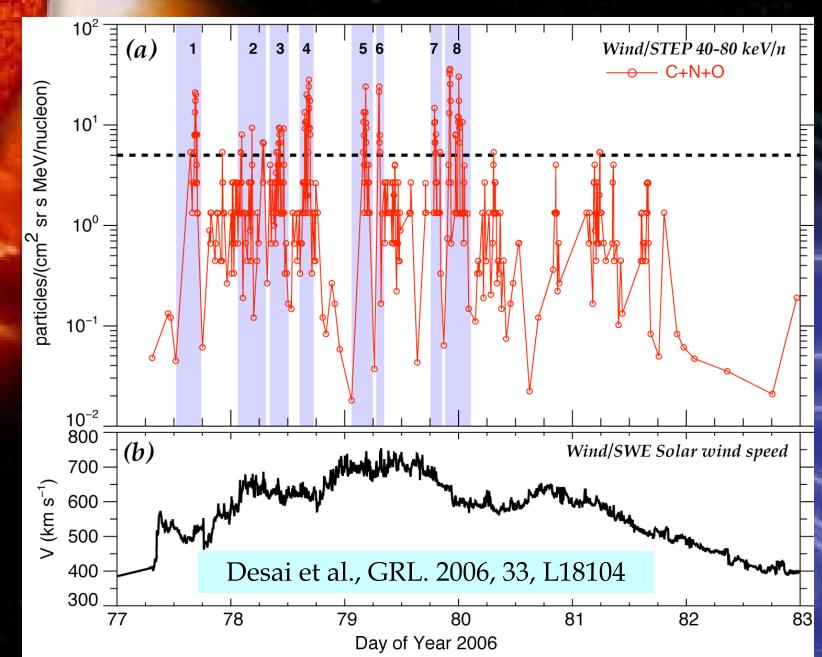


Upstream Events outside Foreshock

- Short duration (~minutes to hours) bursts
- Highly controversial origin and acceleration:

Property	Acceleration at Bow Shock	Acceleration inside magnetosphere + leakage
Composition	Solar wind-like: C, Ne-S, Fe	Ionospheric Ions, O+, N+ & solar-wind-like
Spectra	Exponential in E/Q; cut-off ~150 keV/q - Diffuse ions Reflected ions, ion Beams	Power-laws, extending up to ~2 MeV
Electrons	Only at quasi-perpendicular portions via shock drift process	Up to ~300 keV
Anisotropy	Isotropic, gyrating, beam-like	Free streaming

Upstream Events outside foreshock



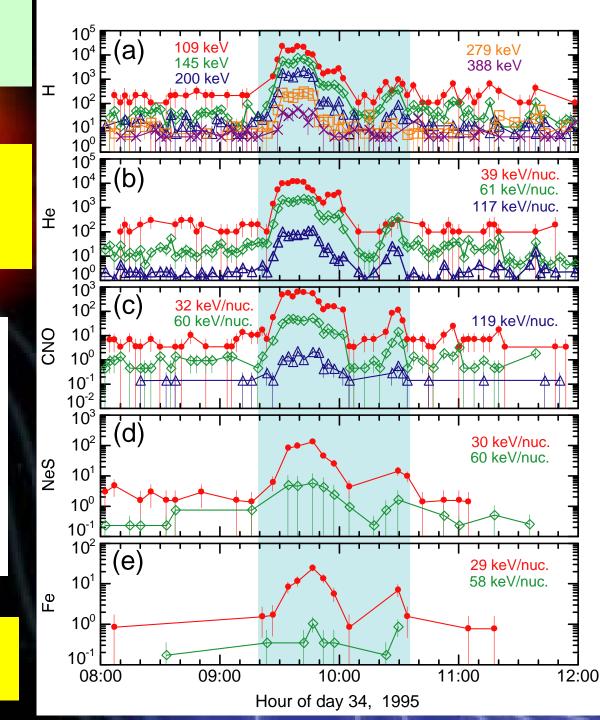
UPSTREAM EVENTS >30 Re

Wind/STEP Timeintensity profiles for 1 event

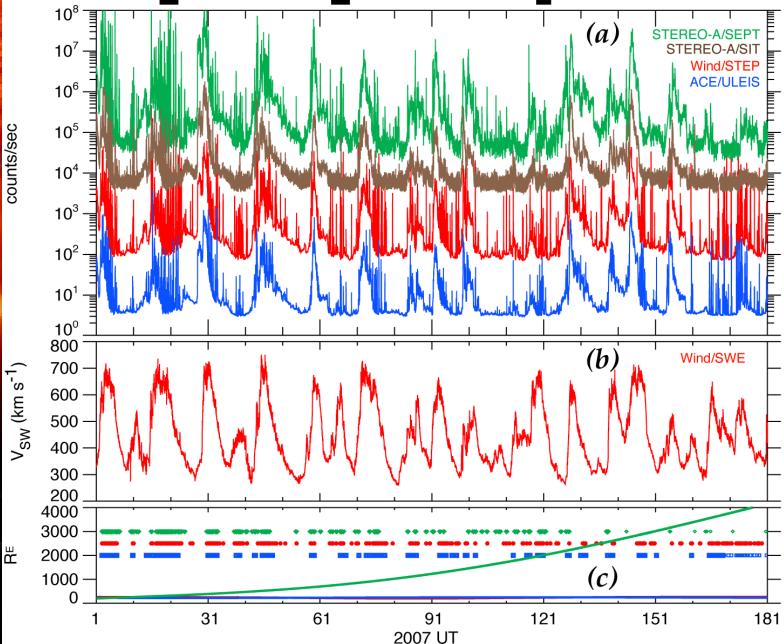
Abrupt intensity increases lasting an hour

Solar-wind species like Ne-S and Fe are present

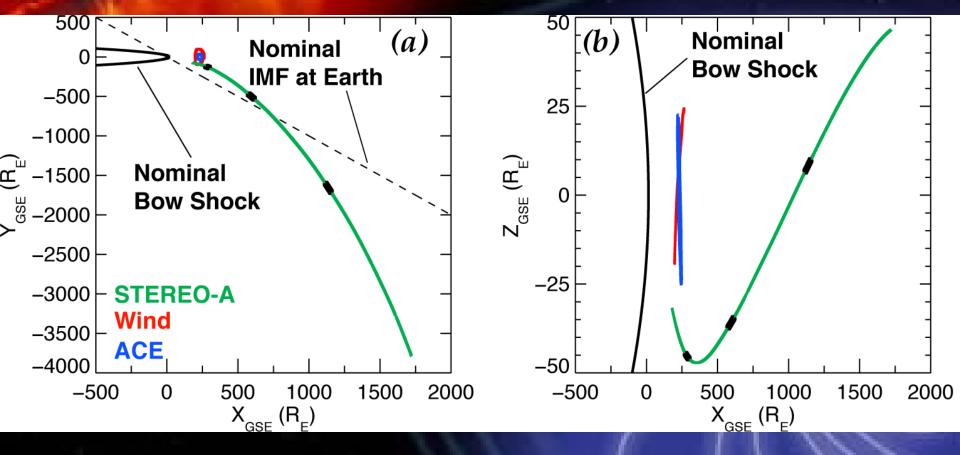
Desai et al., 2000, JGR, vol. 105, pp. 61-78



Overview



ACE, Wind, and STEREO-A during 2007, 1 - 2007, 181

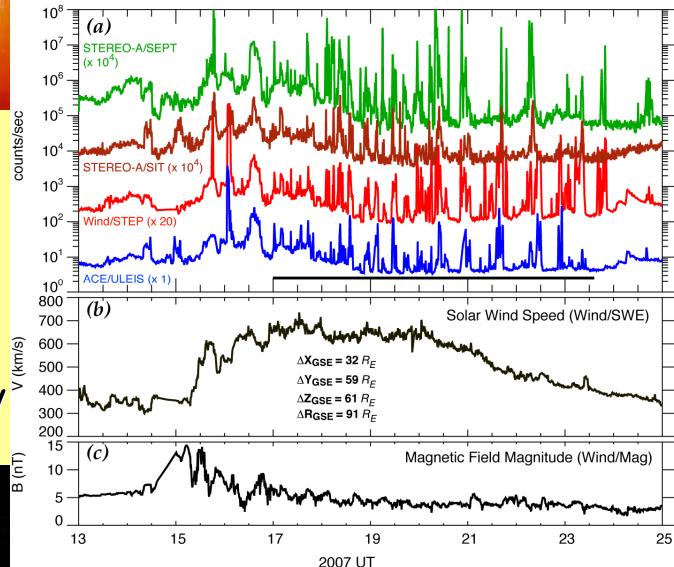


Event Selection and Data Analysis

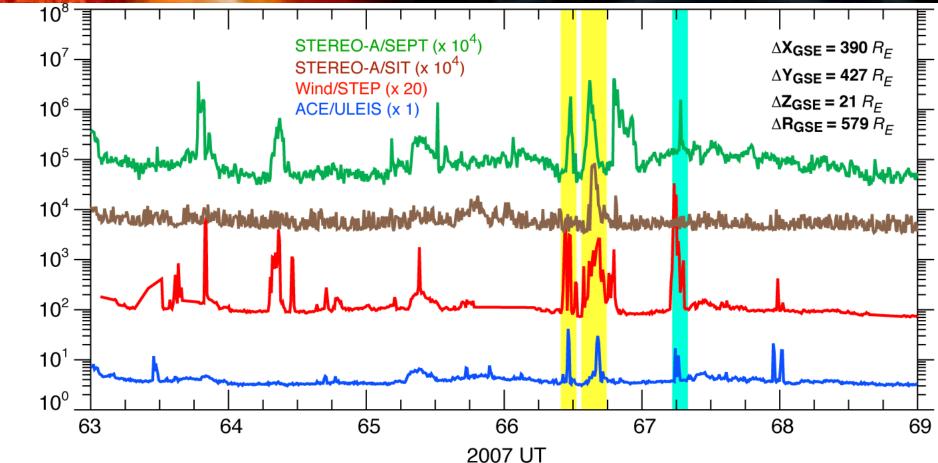
- Use ACE/ULEIS, Wind/STEP, STEREO-A/SEPT and STEREO-A/SIT measurements from 2007, day 1 - day 181.
- Identified upstream events independently at each spacecraft
 - → 300 at Wind, 201 at ACE, and 181 at STEREO-A
- Identified simultaneous events: Events that occurred within 2-hour intervals of ongoing events at Wind
 - → 90 simultaneous events at WIND/STEREO-A

12-day period, showing a sequence of upstream events

Events occur when the S/C are immersed in the high speed solar wind flow after the compression region has passed Earth

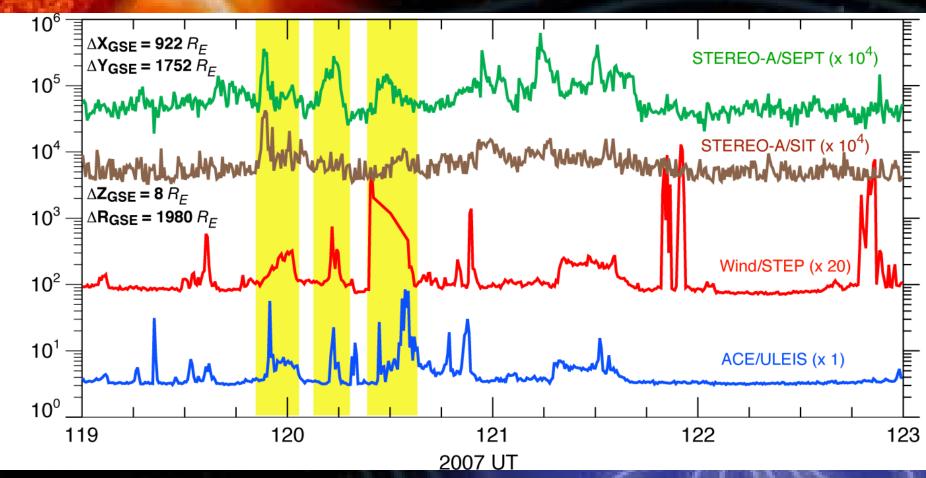


Simultaneous Upstream Events at separation distance of ~580 R_E



counts/sec

Simultaneous Upstream Events at separation distance of ~2000 R_E



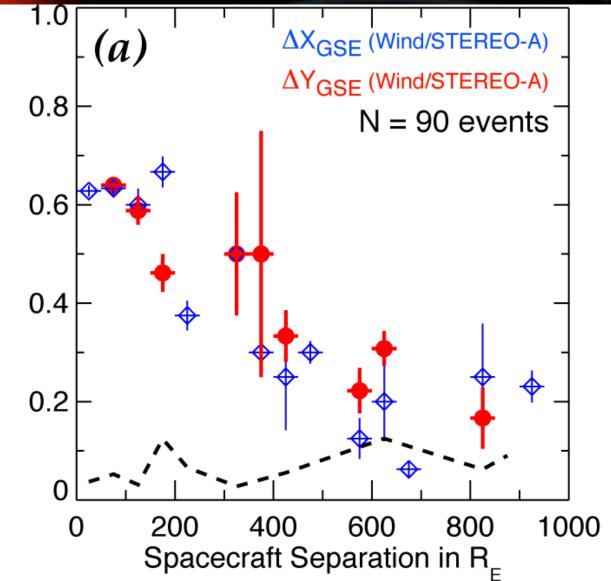
counts/sec

Occurrence Probability of simultaneous upstream events

Probability

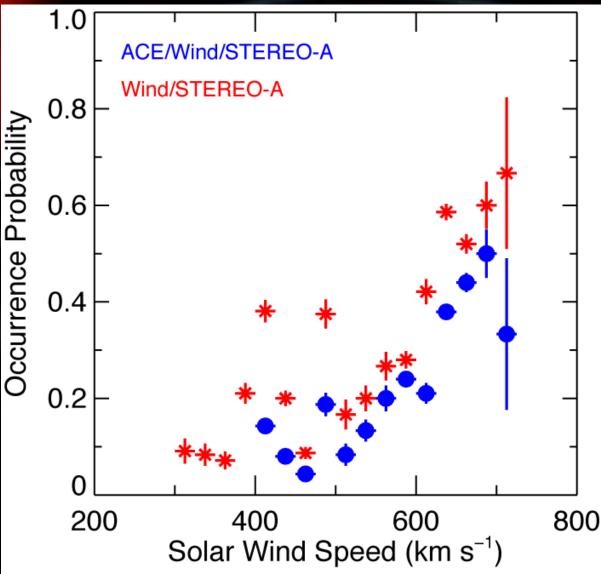
Occurrrence

Upstream events are observed simultaneously at Wind and STEREO-A when the separation distance >800 R_E

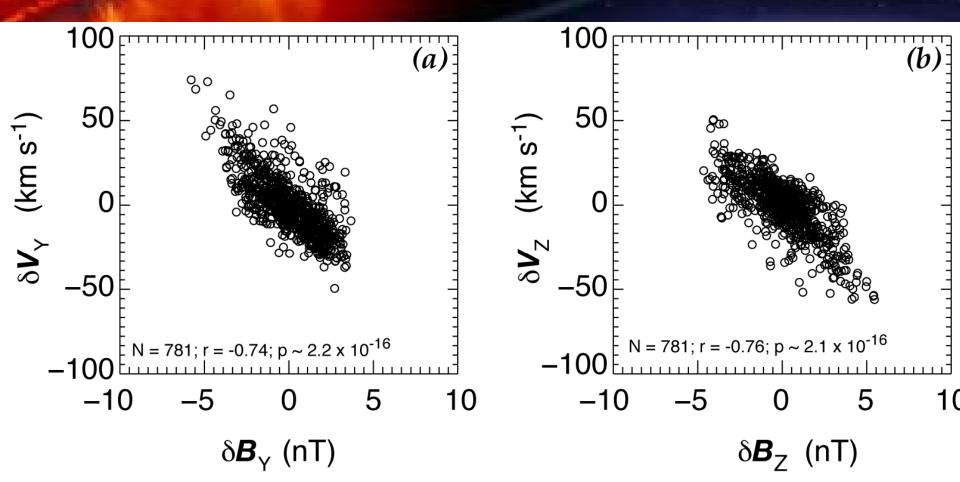


Occurrence Probability of simultaneous vs. *V_{sw}*

Occurrence probability of simultaneous upstream events increases dramatically when solar wind speed increases above ~600 km/s



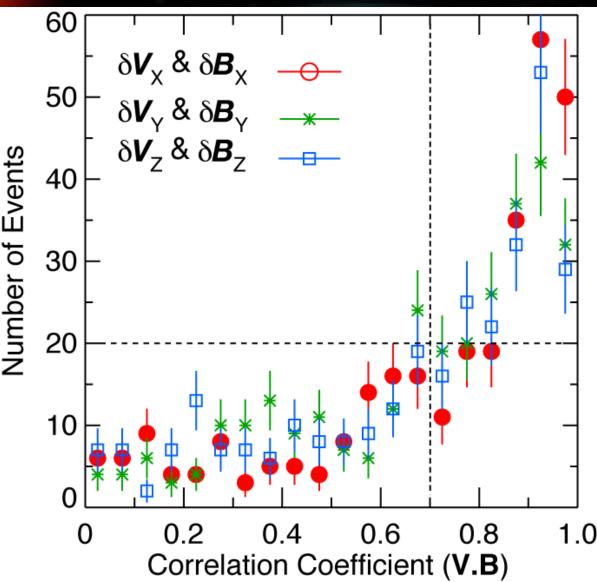
Simultaneous upstream events are accompanied by large amplitude anti-Sunward Alfvén Waves



Number of Upstream events vs V.B correlation Alfvén Waves

Majority (~92%) of upstream events occur in association with outward propagating Alfvén waves

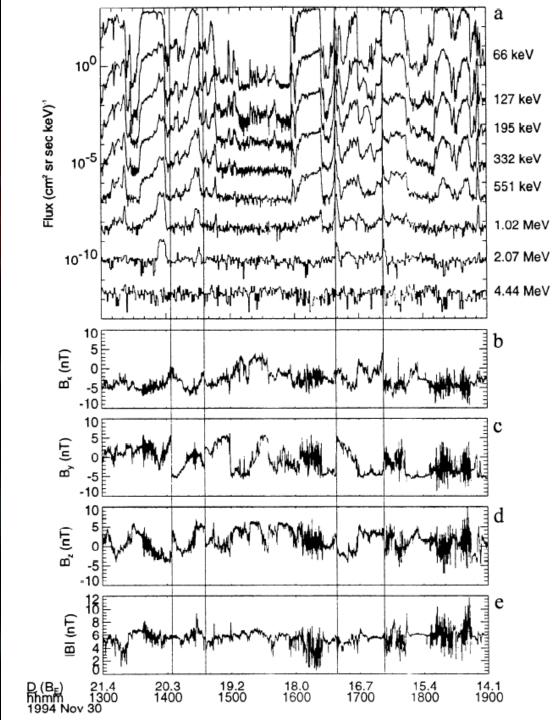




Freeman & Parks (2000)

Upstream ions are accelerated by reflection between two converging magnetic mirrors: (1) large IMF rotations and (2) the Earth's bow shock

Our Results show that these "IMF rotations" are large amplitude Alfven waves that are embedded in high-speed solar wind streams



Summary

- STEREO-A observed upstream events even when it was separated from Earth by ~1750 R_E in the radial and ~3800 R_E in the lateral directions
- Occurrence probability for measuring simultaneous upstream events at L1 and STEREO-A remained high (~20-30%) even at large radial and lateral separations (~800 R_E)
- Occurrence probability of simultaneous upstream events at L1 and STEREO-A increases with solar wind speed and in association with anti-sunward propagating large amplitude Alfvén waves.

Conclusions

- Upstream ion events originate from a global source region that is comparable to the size of the Earth's bow shock
- The presence of large amplitude Alfvén waves embedded in high-speed solar wind streams facilitates the scatter-free propagation of upstream ions in the regions traversed by STEREO-A
- The Alfvén waves could also play a critical role in accelerating solar wind and suprathermal ions
 - → Following the passage of the compression region past the Earth, the expanding bow shock could act as converging magnetic mirrors and accelerate by reflection and first-order Fermi processes (e.g., Freeman & Parks 2000).