

Capabilities of coordinated white-light and radio wave observations

Stereo meeting, Hamburg, 1-4 May 2005

M. Pick,

A new web site for radio monitoring

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R. Romagnan and A. Bouteille
LESIA, Observatoire de Paris

Why a web site for radio monitoring ?

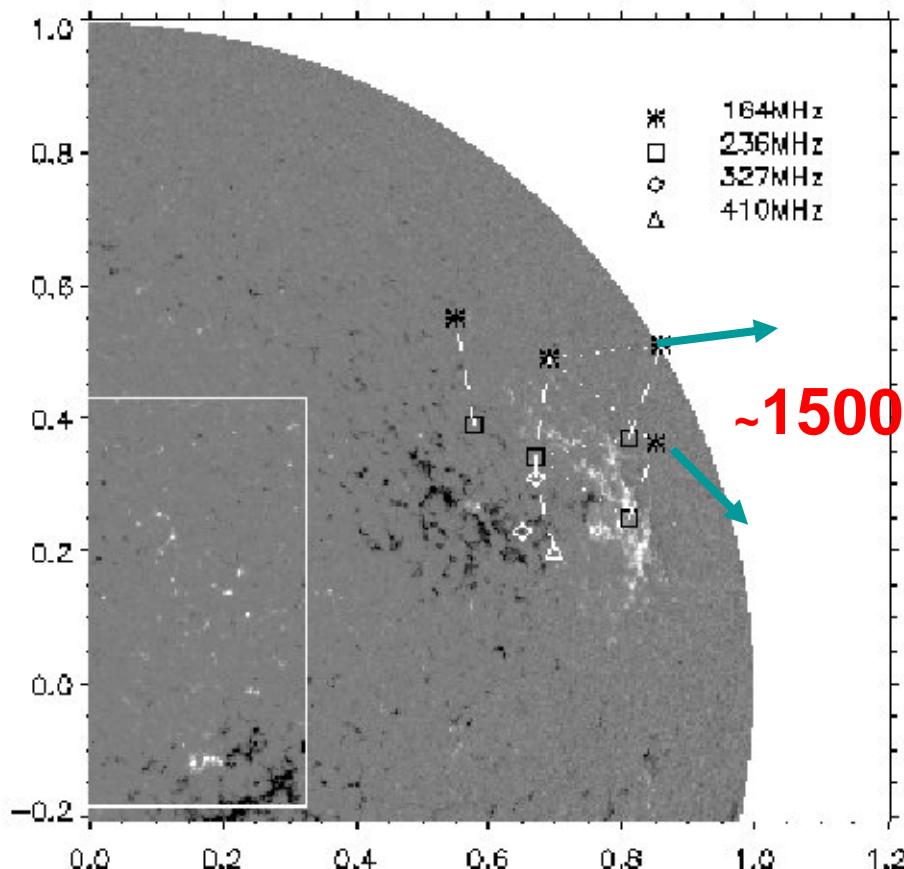
Main objectives

- Radio associated with CMEs, onset, development
- Electron beams from the low corona to the interplanetary medium

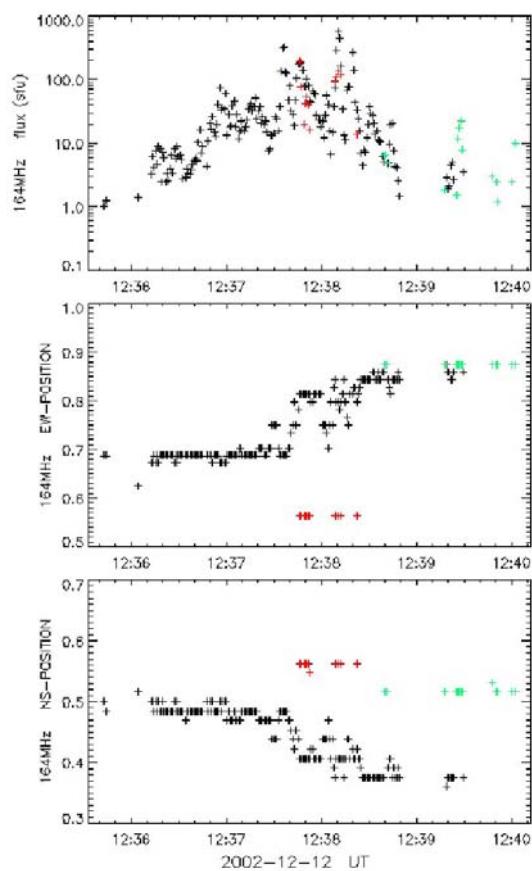
Web site under construction

- will not include microwave data
- will include dm-km data
- **SECCHI CME summary** (R. Howard, A. Vourlidas)
- links with other web sites

Fast narrow CME 12 December 02

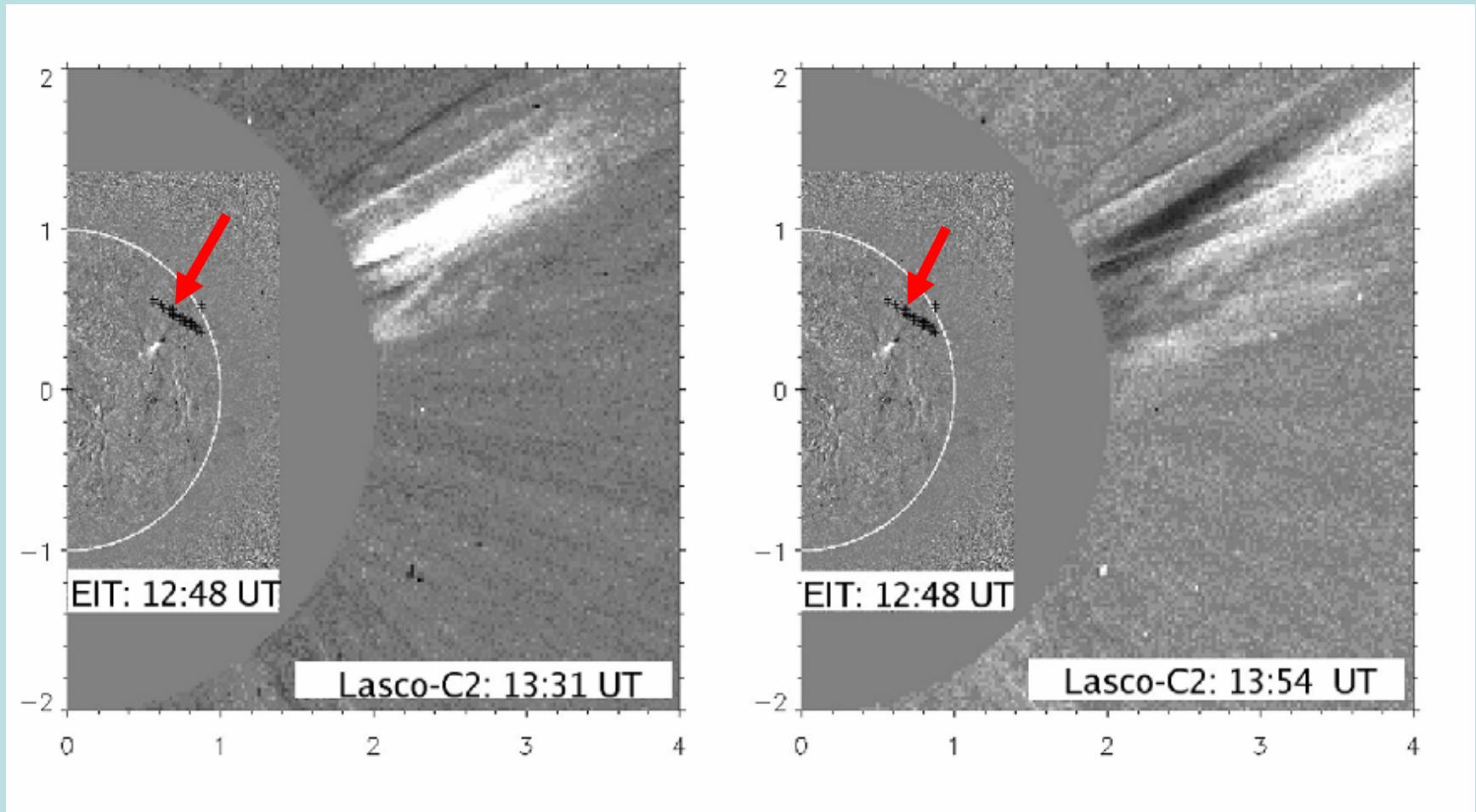


12 Dec, 2002 UT



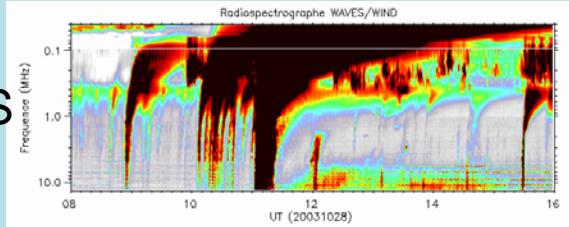
NRH 12:36-12:40 UT
Cadence 10s, 1s
Solar soft

12 December 2002

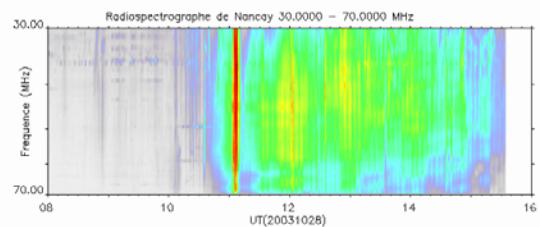


28 October 2003

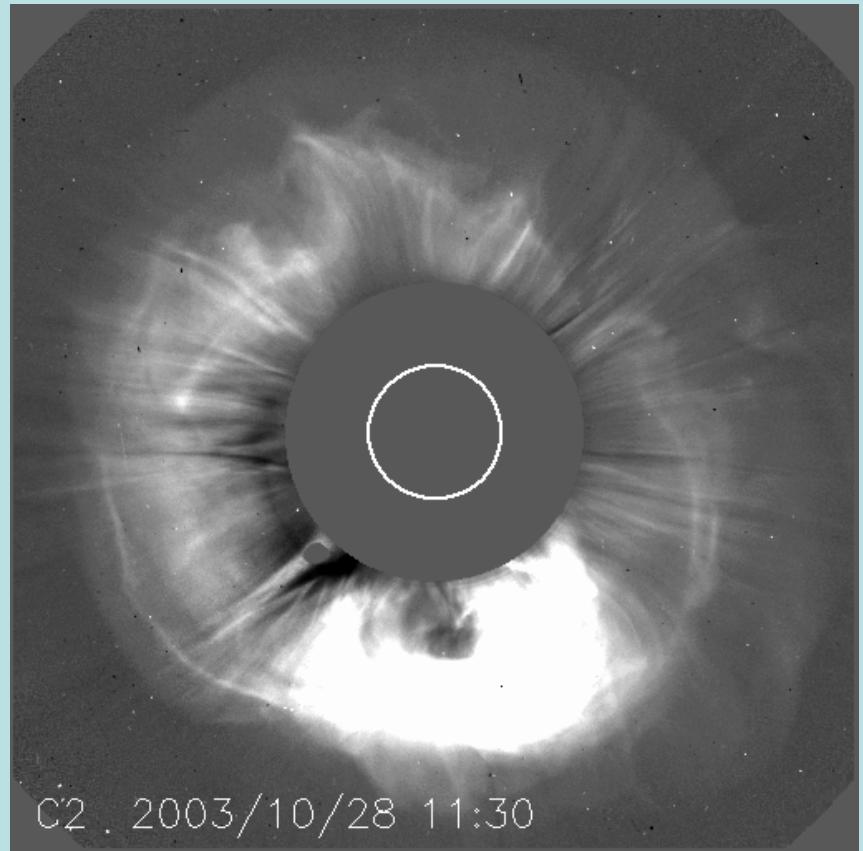
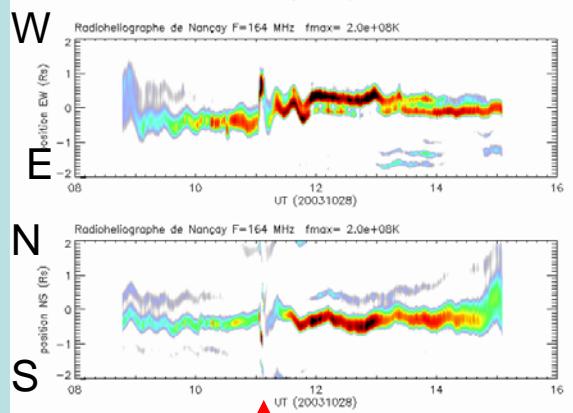
WAVES



DAM



NRH



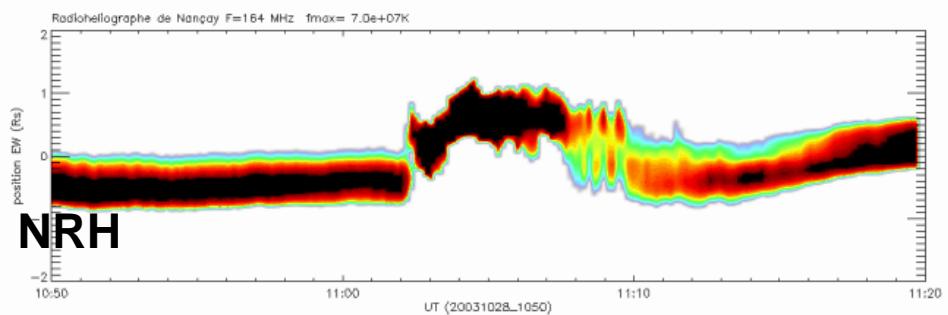
Fast wide CMEs On-Disk development

28 October 2003

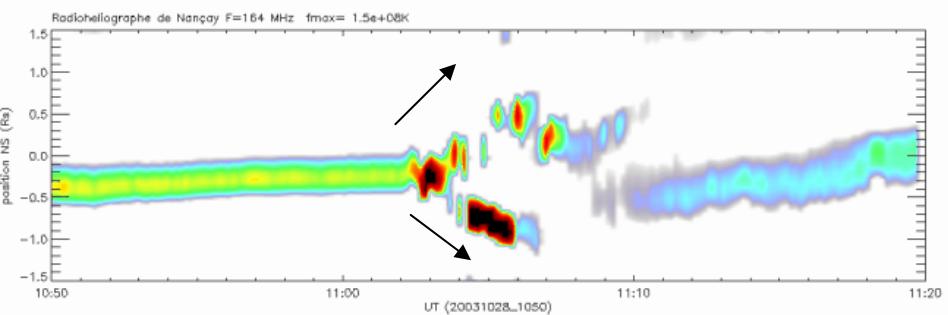
W



E



N

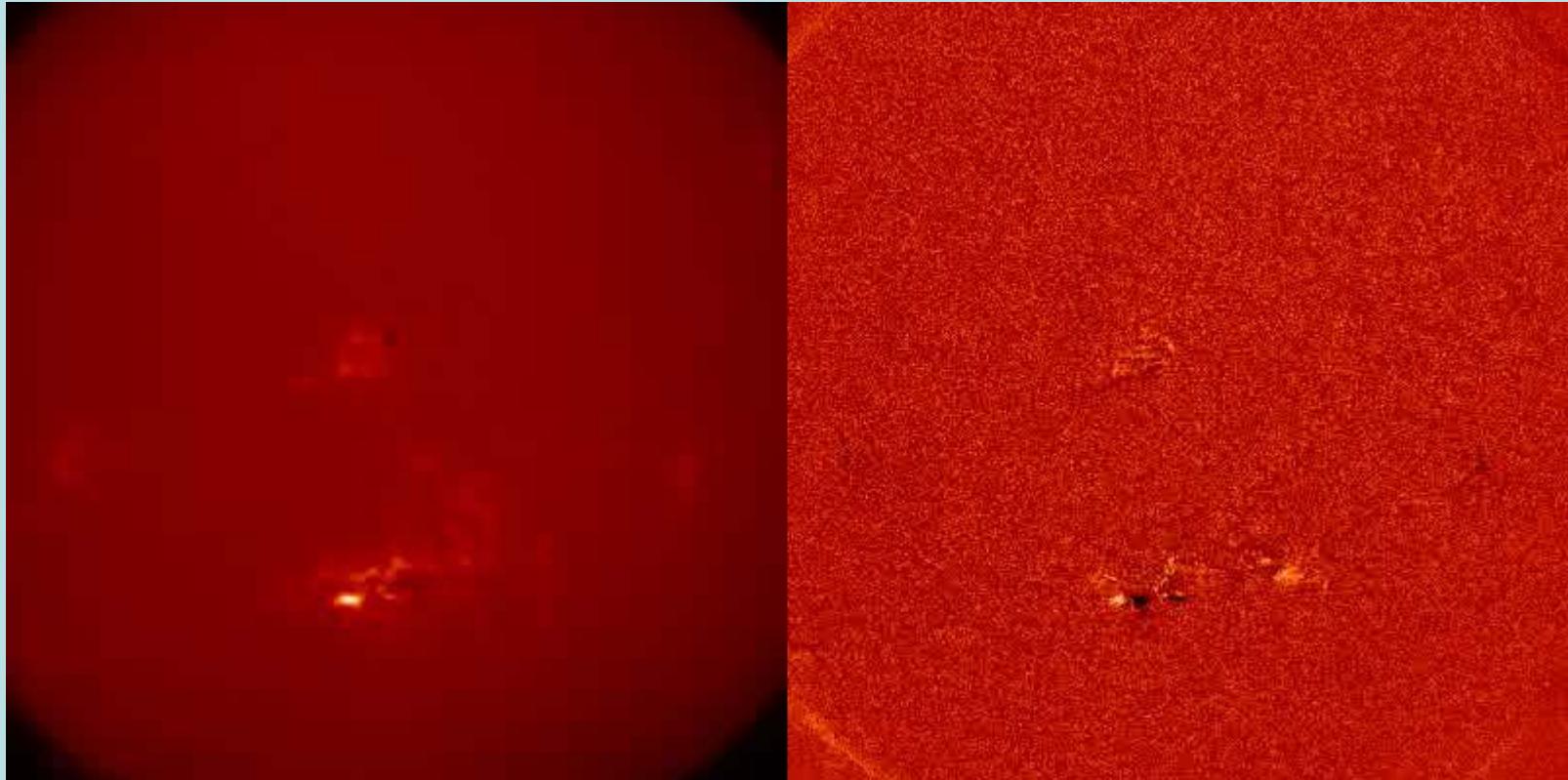


S

Radio sources **spread, suggest a disturbance V ~2500 km/s**
24 events (NRH observing time 1996-2000) spreading
(May 02 1998, association with Moreton wave,
Pojholainen et al., 2001, ApJ)

Heliograph H α « λ variable » de Meudon

Cadence 1 mn with 3 λ

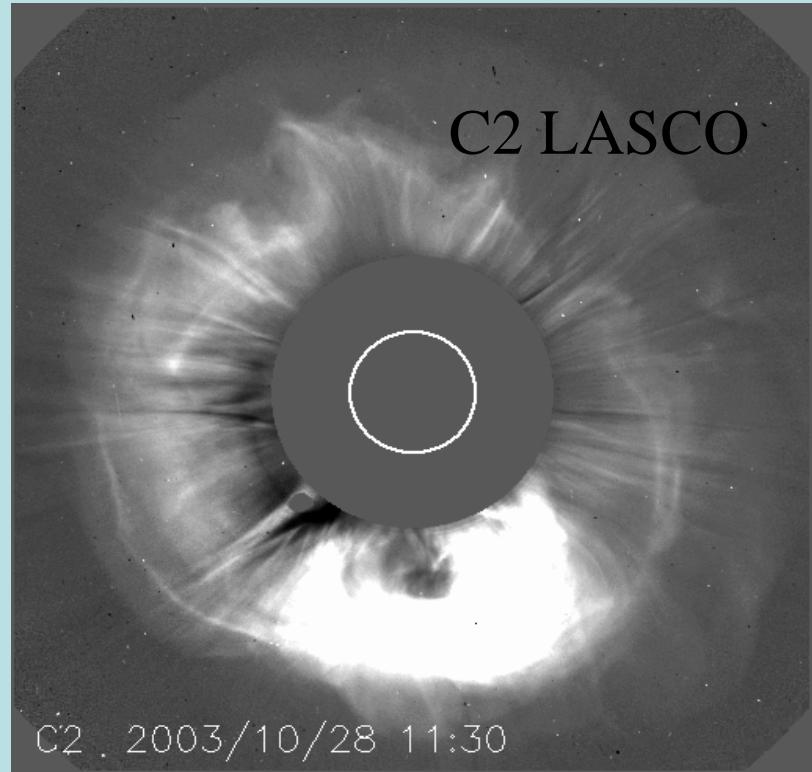
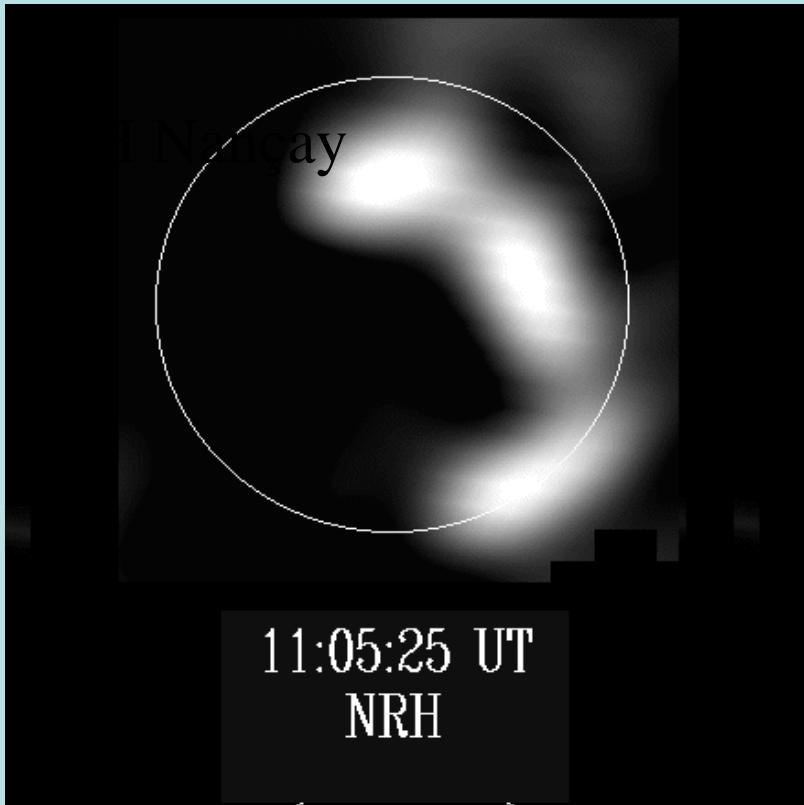


Intensity center of line

Running differences

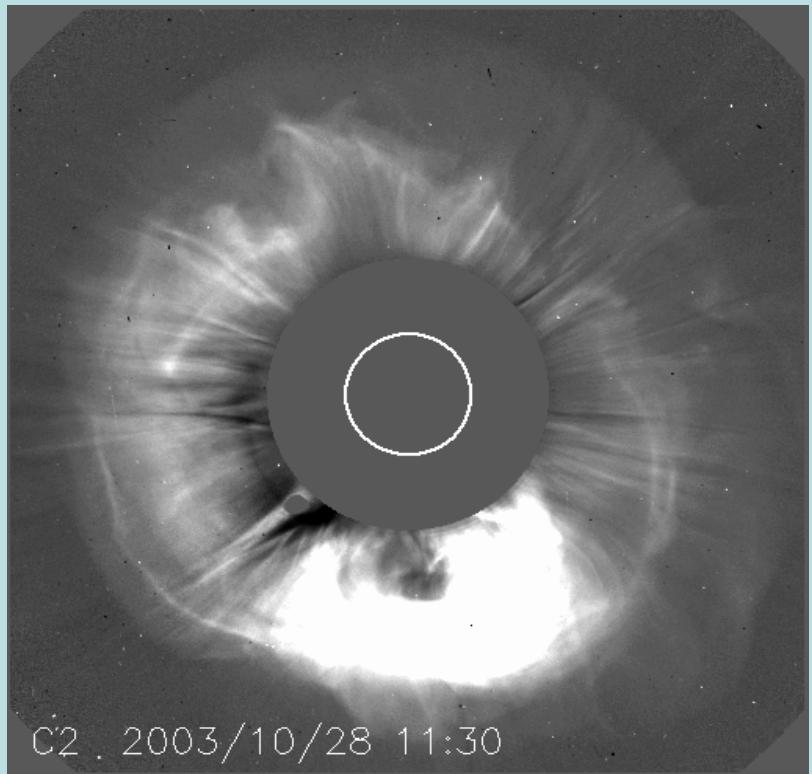
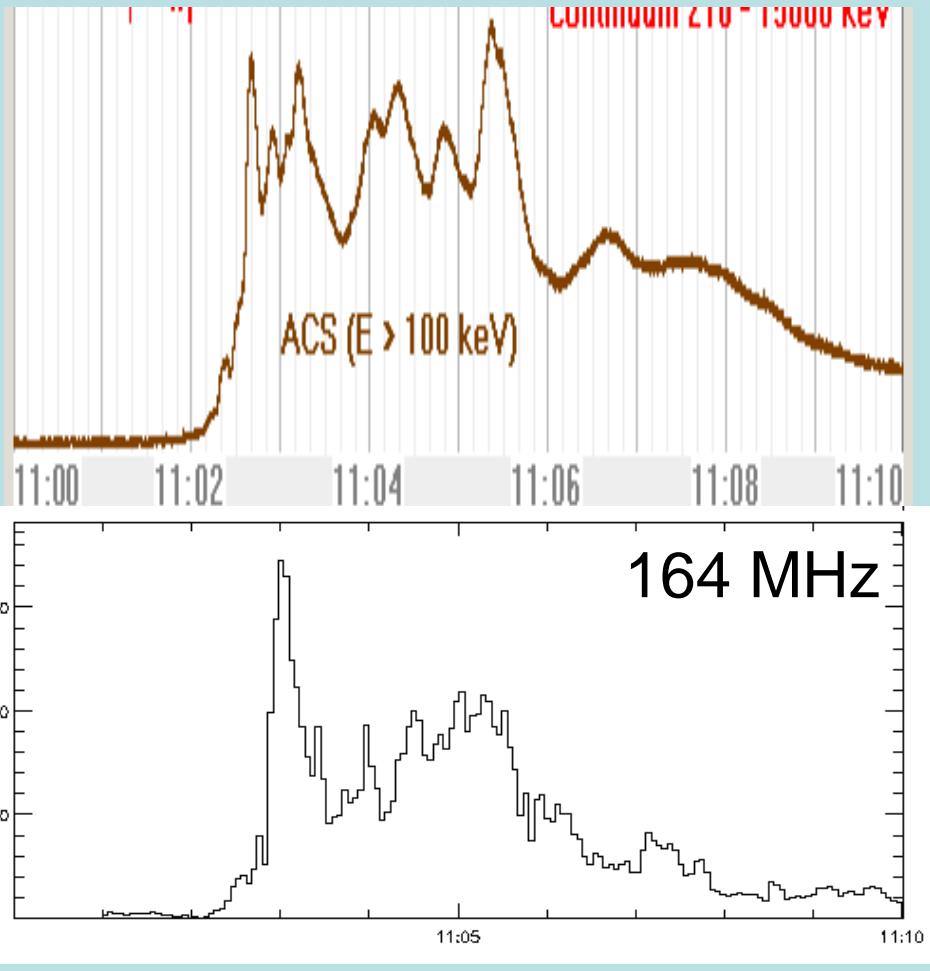
28 Octobre 2003

October 28 2003 Onde de Moreton et CMEs

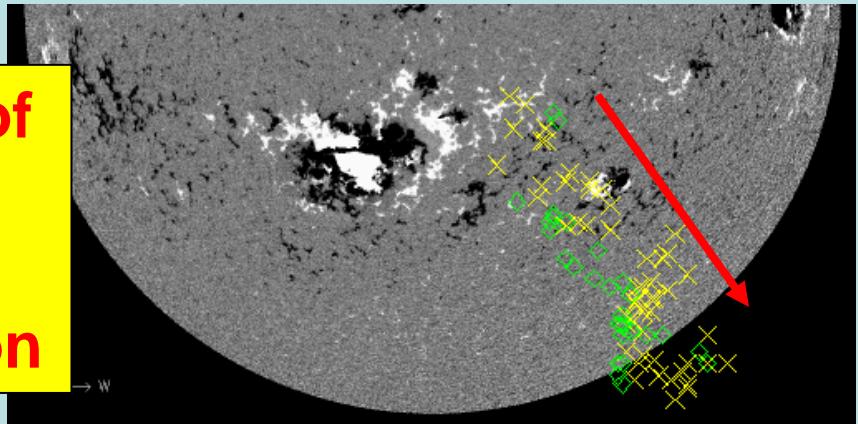


Radio imaging at 164 Mhz

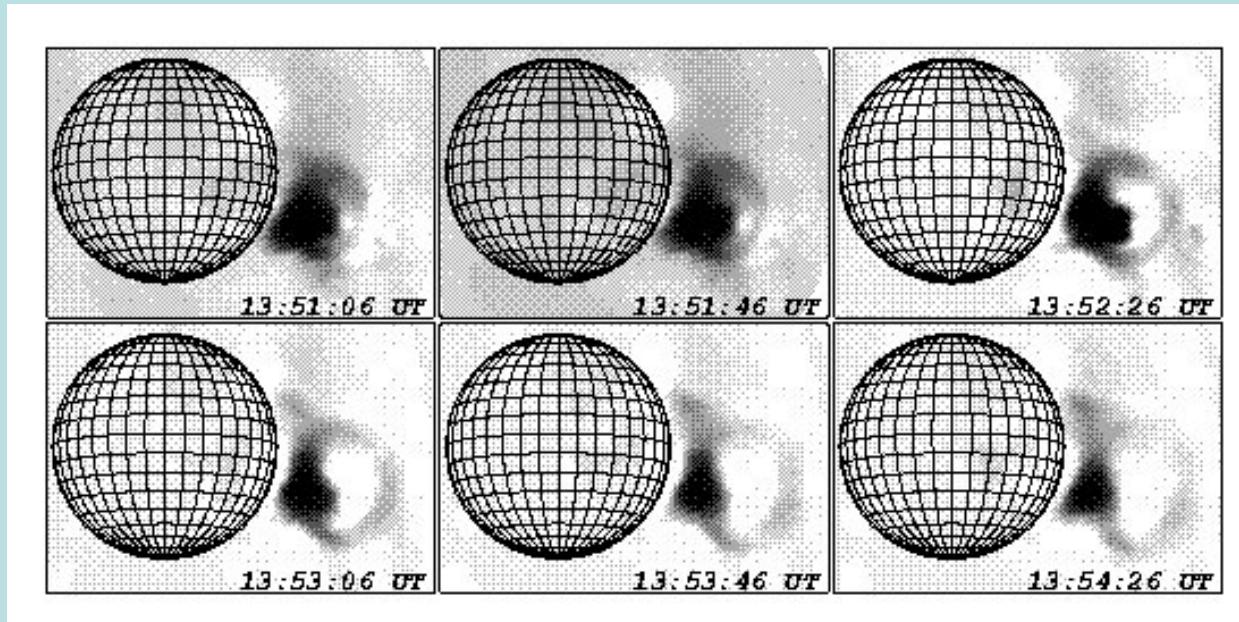
- traces the CME development and the regions of B interaction
- The coronal wave and the CME encompasses most of the solar disk



- Electrons trace the expansion of the arch system (~ 2500 km/sec)
- Coronal wave develops along flank of CME In lateral expansion



Radio CMEs



Web page

Concluding remarks

- Weekly deliver
 - NRH clean images
 - Image 10 ko
- Work
 - S-Waves
 - Other radio spectrographs ?
 - Survey (composite) 1997-to-day