PROBA2: READY FOR SCIENCE & SPACE WEATHER

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9th SECCHI meeting * Trinity College Dublin, Ireland * March 22, 2010







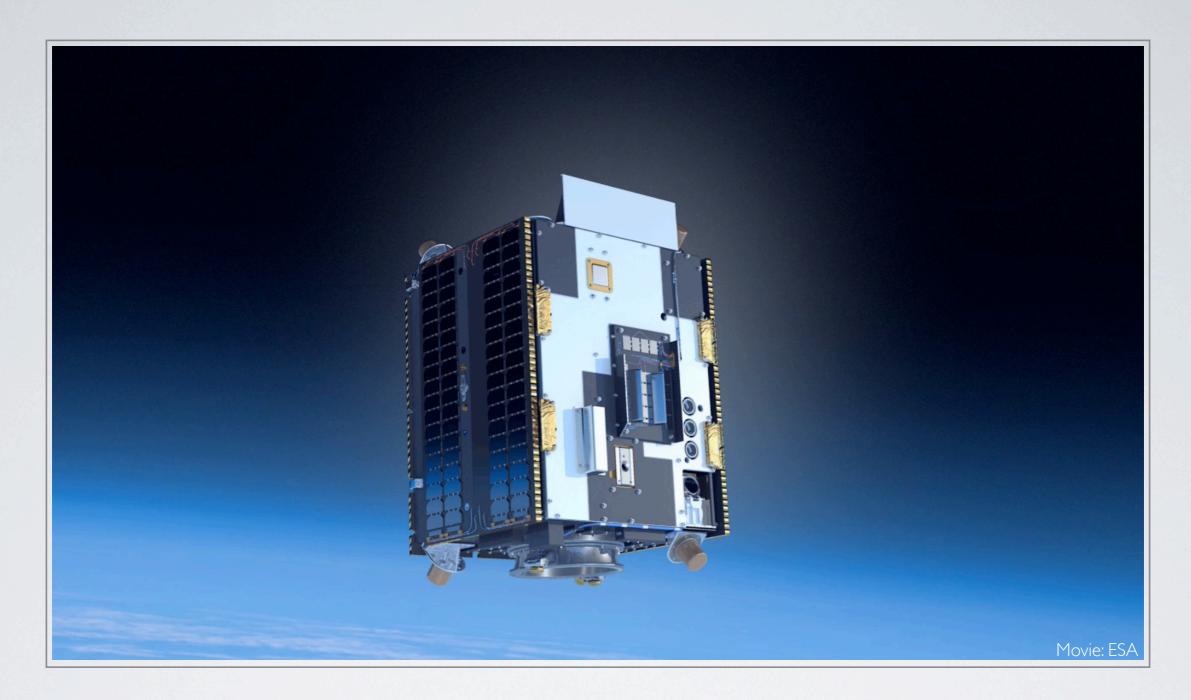






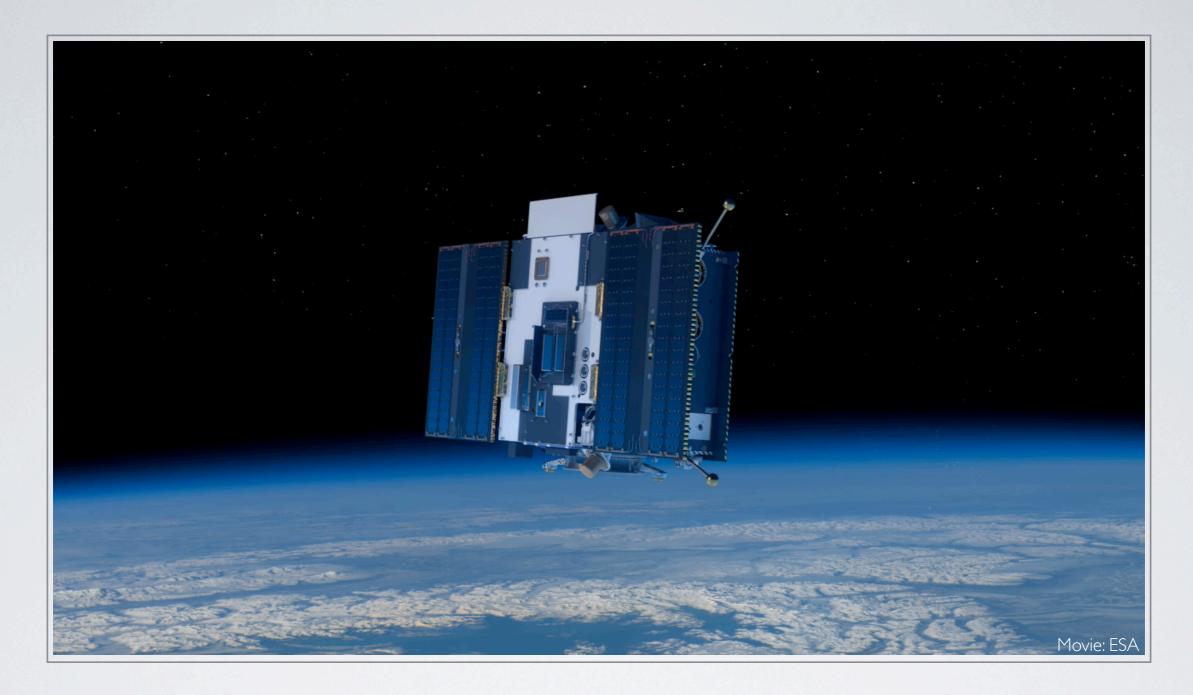
PROBA2

Microsatellite in sun-synchronous orbit Launched on November 2, 2009 725 km altitude



ESATECHNOLOGY MISSION

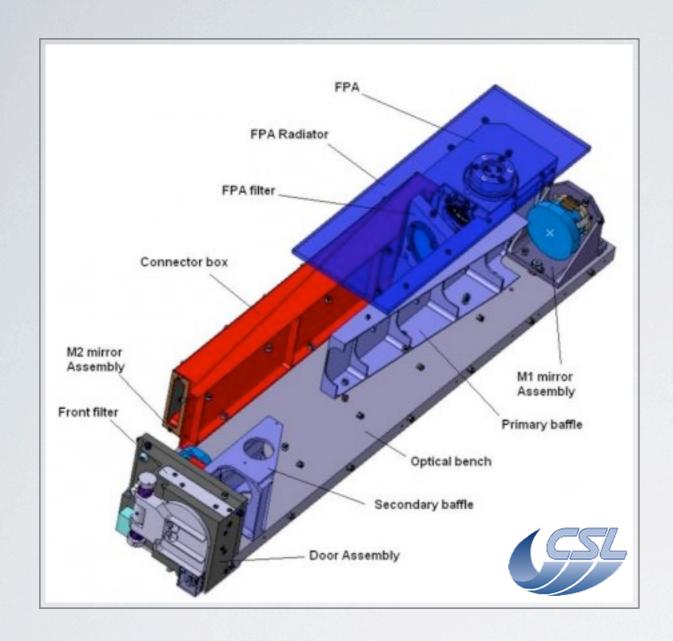
4 innovative instruments: SWAP, LYRA, TPMU, DSLP 17 technological experiments made in Belgium

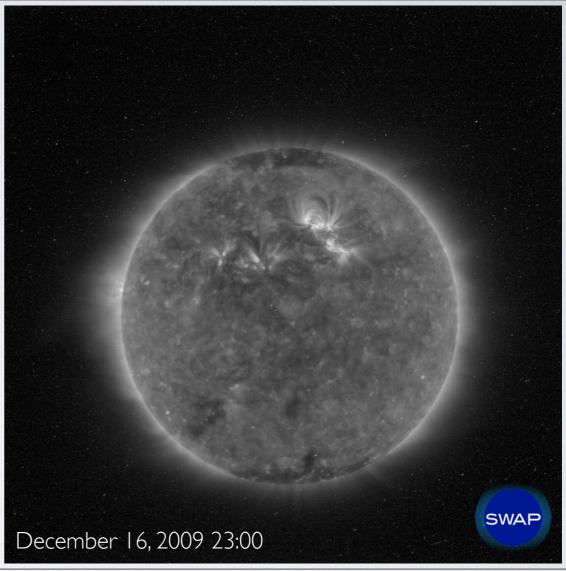


ESA SCIENCE MISSION

SWAP and LYRA observe the Sun in UV and X rays commanding and data processing at P2SC (ROB, Brussels) PI SWAP: D. Berghmans PI LYRA: J.F. Hochedez

First RAW image with rough centre pointing



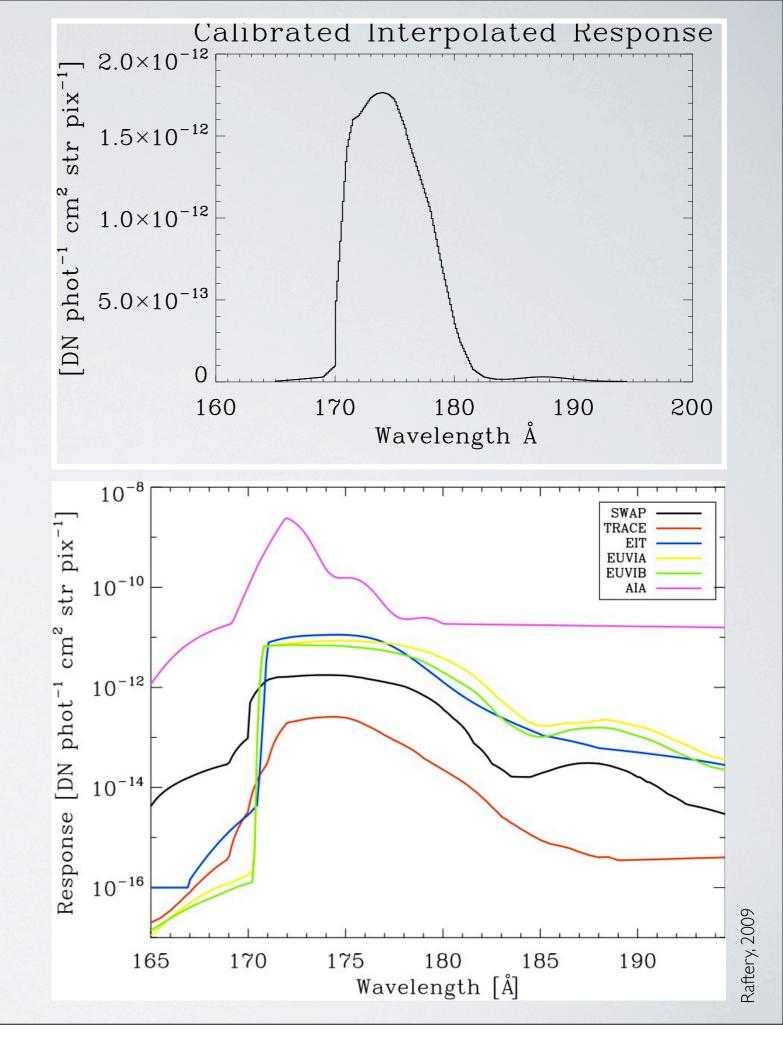


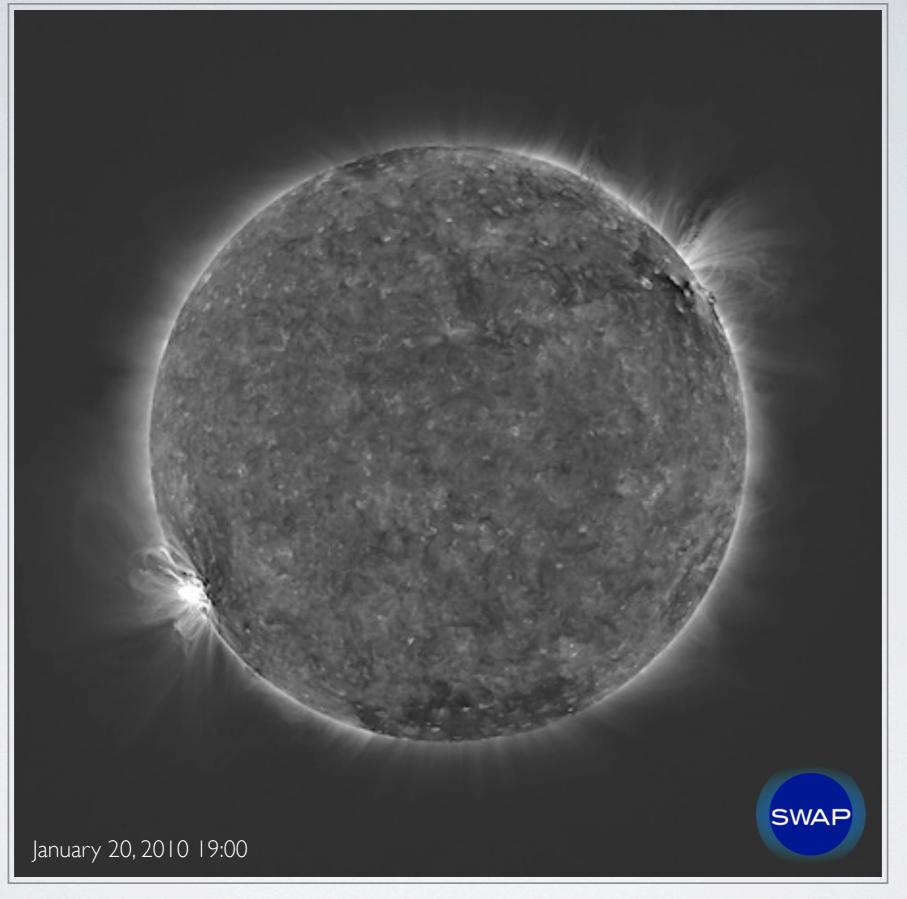
SWAP EUV IMAGER

Exercise in miniaturisation: off-axis Ritchey-Chrétien scheme Observes the 1 million degree corona in EUV light First light on December 14, 2009 Nominal operations start NOW

SPECTRAL RESPONSE

Peak at 17.4nm Comparable to EUVI





PROCESSED IMAGE

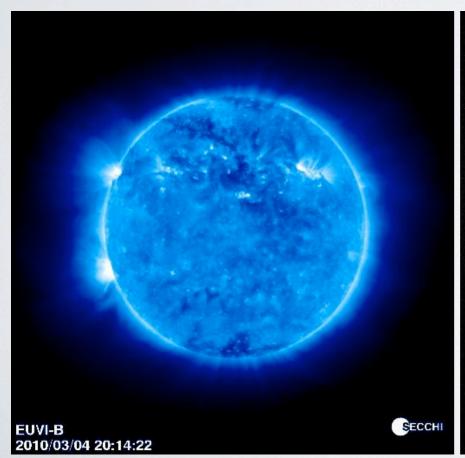
SWAP vs. EUVI

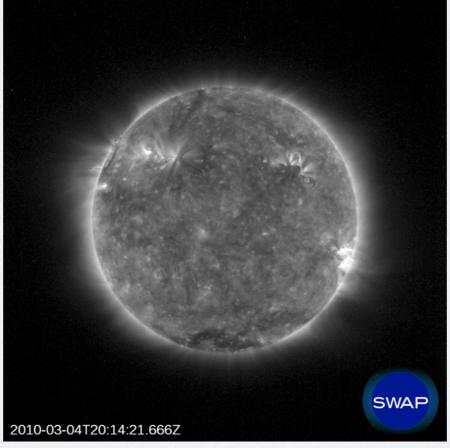
EUVI: 17.1, 19.5, 28.4, 30.4 nm

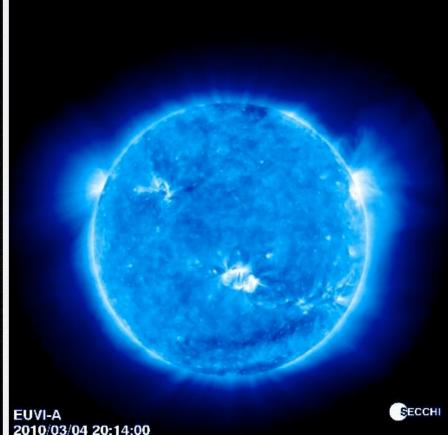
- CCD detector with shutter
- cadence: 2h (171)-5' (195)
- 2048×2048 pixels
- stereo view: ahead & behind Earth
- sun-centered FOV

SWAP: 17.4 nm

- shutterless CMOS detector
- ≈ I minute cadence
- 1024×1024 pixels
- inside magnetosphere, Earth view
- flexible off-pointing







SWAP FOR SCIENCE & SPACE WEATHER

- High temporal cadence up to 18s (Imin nominal)
- Onboard data-processing and prioritization
- Data available in near-real time (9 passes/24hrs + priority nr)
- Off-pointing capabilities (even automatically)
- Several compression schemes: LZW + JPEG (+options)
- Flexible commanding from the Science Center

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ANNULAR SOLAR ECLIPSE



January 15, 2010, 06:00 UTC



ACTIVE REGIONS

1st M-flare

January 18-21, 2010



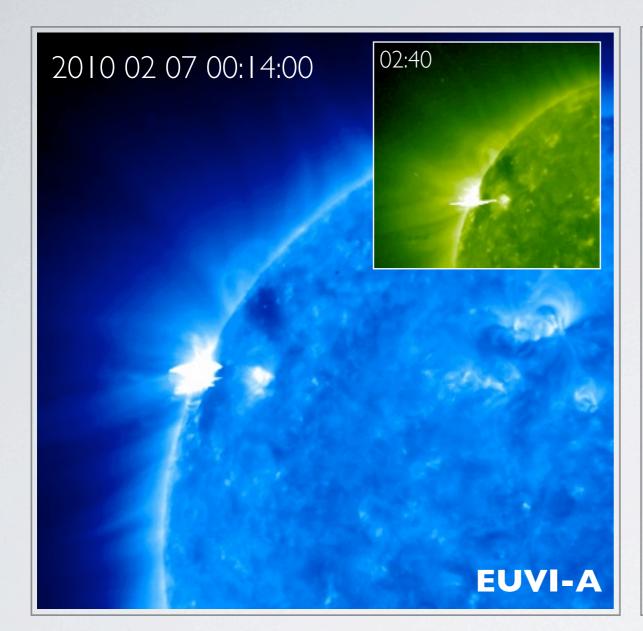
of new

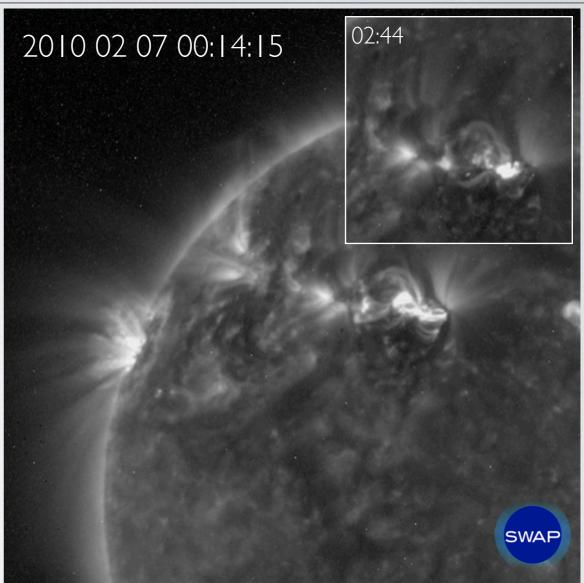
cycle



SOLAR FLARES

limited blooming due to CMOS detector high cadence up to 18s LYRA gives detailed temporal evolution (10ms)





SOLAR FLARES

limited blooming due to CMOS detector high cadence up to 18s

LYRA gives detailed temporal evolution (10ms)

LYRA FLARES

4 channels:

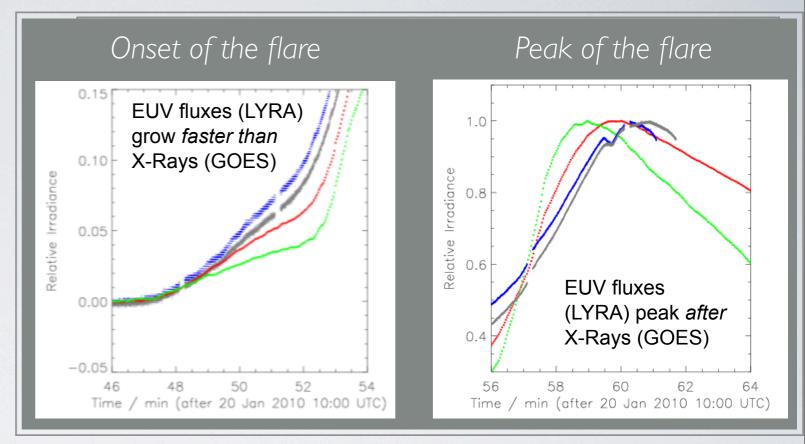
Ly alpha: ~ I 20nm

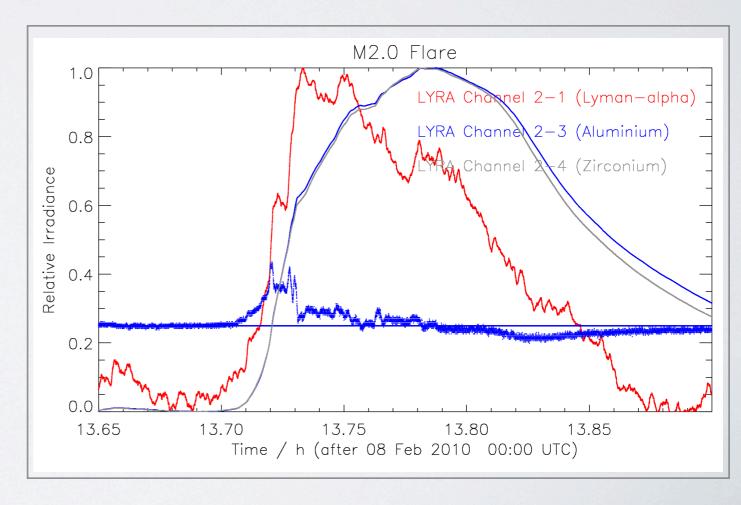
Herzberg: 200-220nm

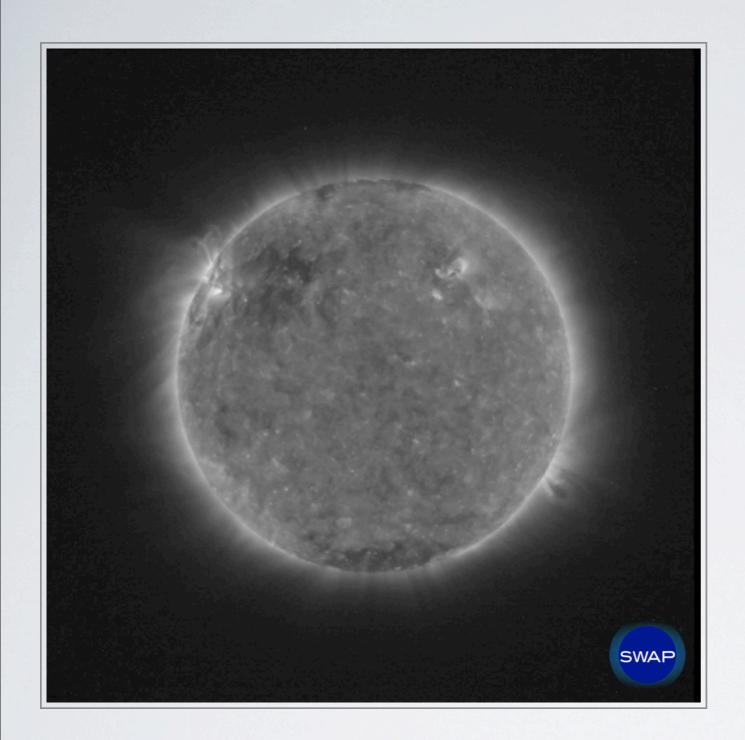
Al: XUV & EUV (incl. Hell)

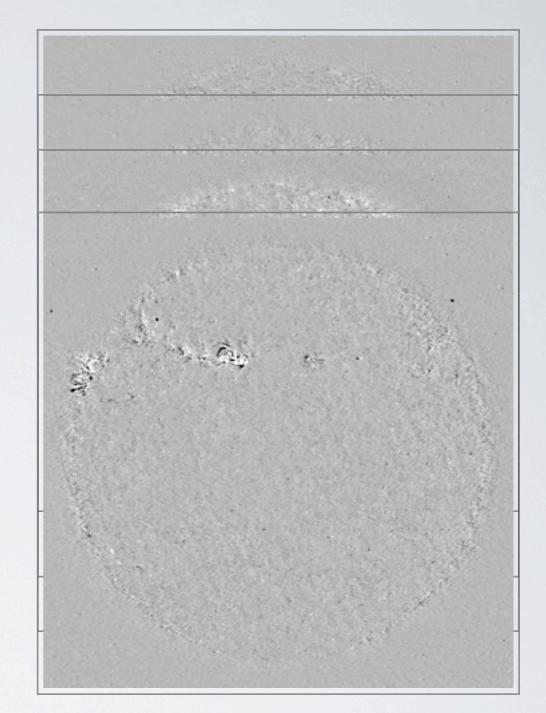
Zr: XUV & EUV(excl. Hell)

high temporal resolution up to 10ms



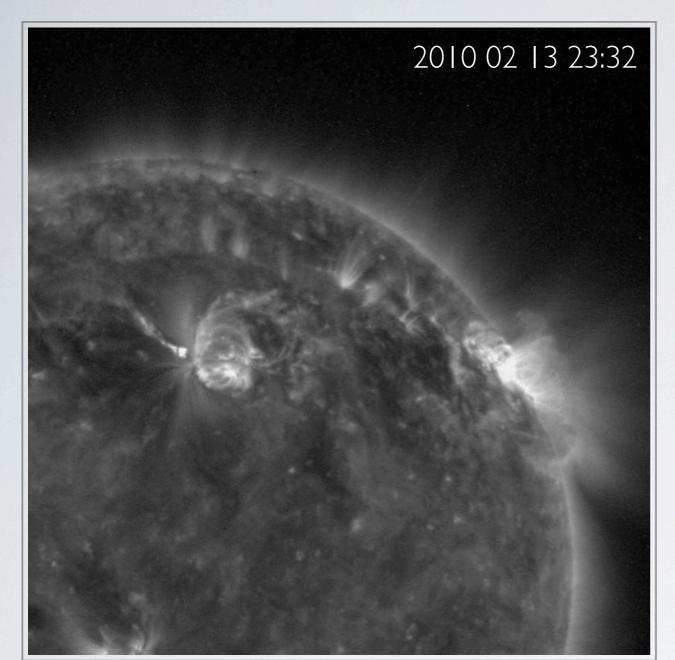


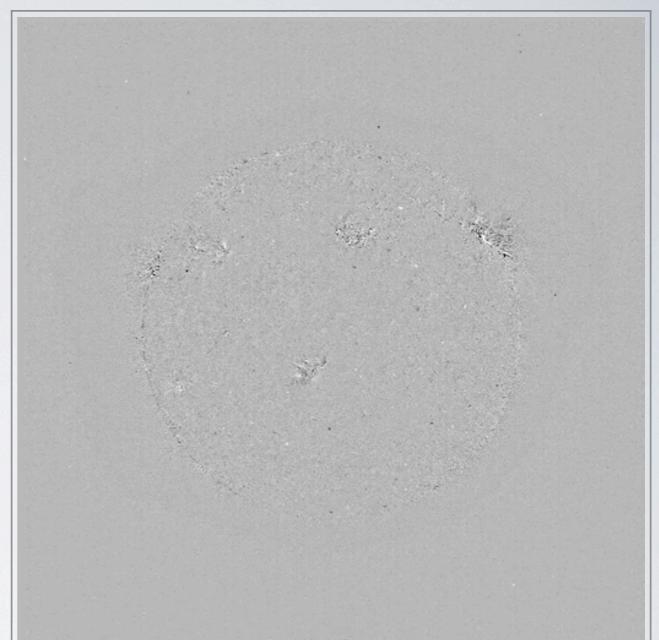




EIT WAVES

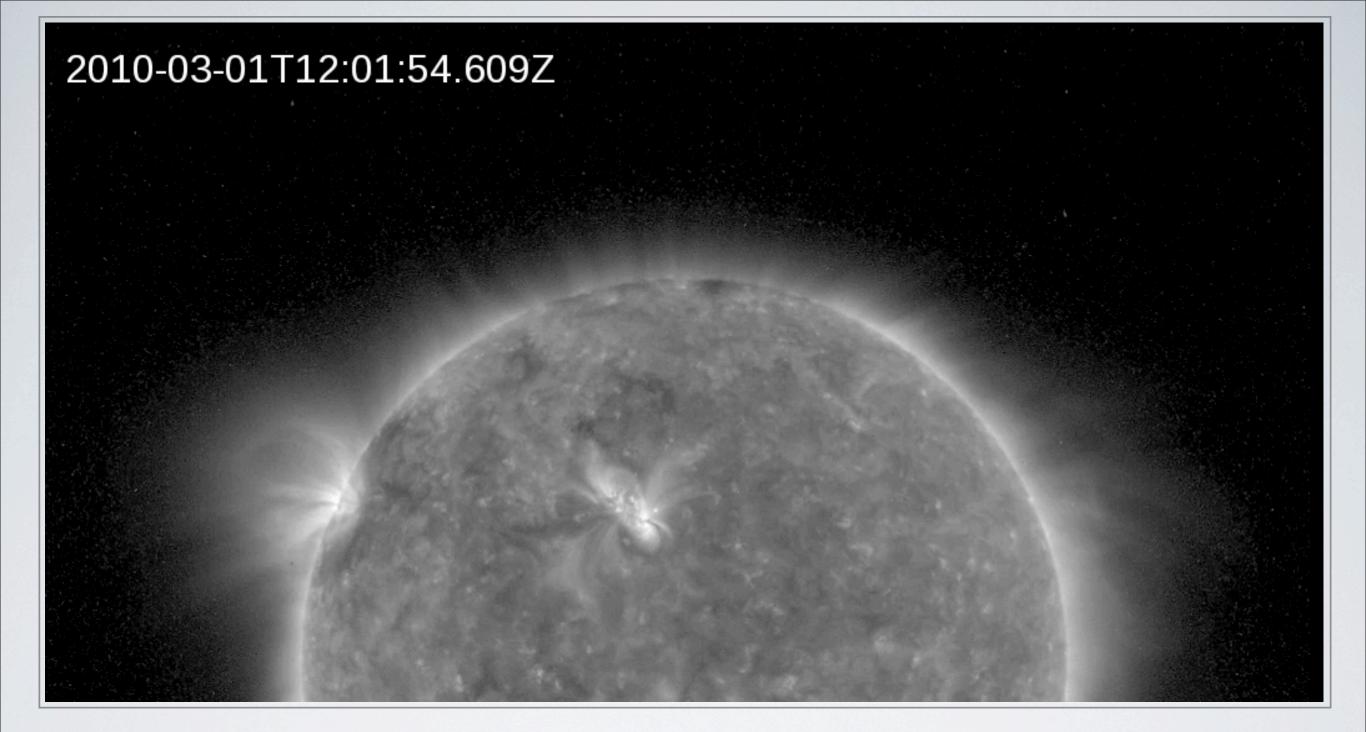
several waves associated with the flares of Feb. 5-6, 2010





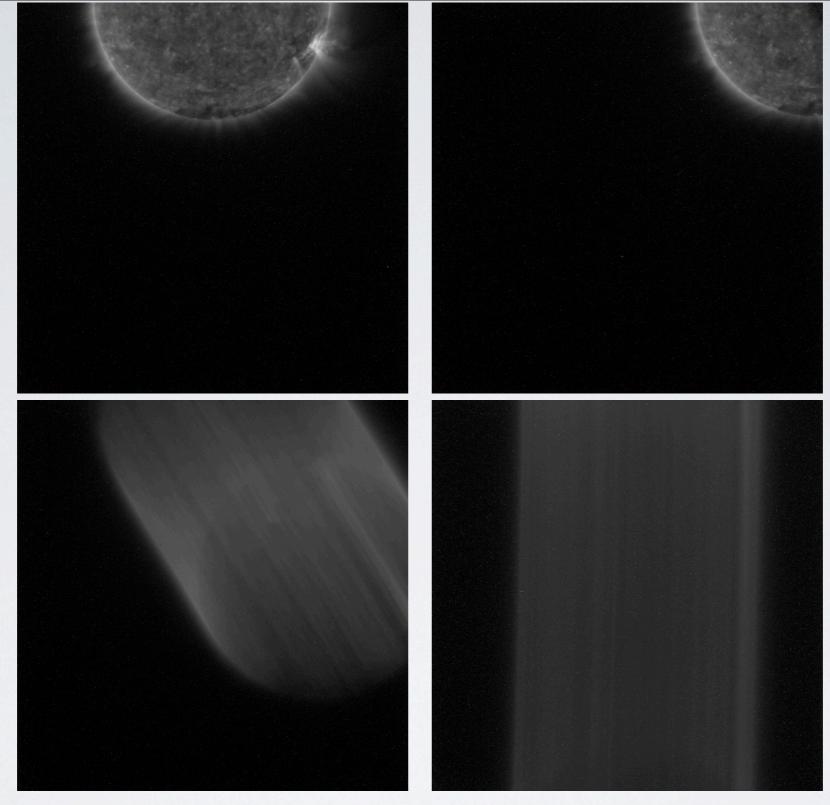
EUV JETS

12-13 February: co-temporal jets leave from 2 separated ARs speed roughly 500km/h observed both in SWAP and <u>EUVI</u>



CORONAL MASS EJECTIONS

waiting for more clear events in SWAP off-limb region automatic CME tracking probably hampered by unstable pointing



OFF-POINTING

up to now mostly interesting for calibration could be used for CME tracking

OPEN DATA POLICY

Data will be freely available to all users from mid April onwards

http://proba2.sidc.be

All data ordered in year/month/day folders Fancy data browser to come

Raw Engineering FITS: reformatted, decompressed, long header Base Science Data FITS: (preliminary) calibrated, science header PNG files: for quicklook purposes (available now)

SSW will have software trees SWAP & LYRA thanks to Trinity College Dublin

HOWTO BE INVOLVED?

Scientists are welcome to:

- use PROBA2 data,
- propose special observation campaigns

Guest Investigator Program welcomes proposals for dedicated (joint) observations in the frame of a science project

- funds available for a stay at PROBA2 Science Center
- scientist can take part in the commanding of the instruments
- will gain expertise in the instrumental effects
- Announcement: May 3, 2010
- Gl proposal deadline: June 1st (visits from Sep 2010 onwards)
- First Science Working Team Gl selection: June 7-9

FOR MORE INFORMATION

http://proba2.sidc.be/

