

PICKUP HELIUM OBSERVATIONS WITH STEREO PLASTIC A & B



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WORK IN PROGRESS

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STEREO Pickup He Observations

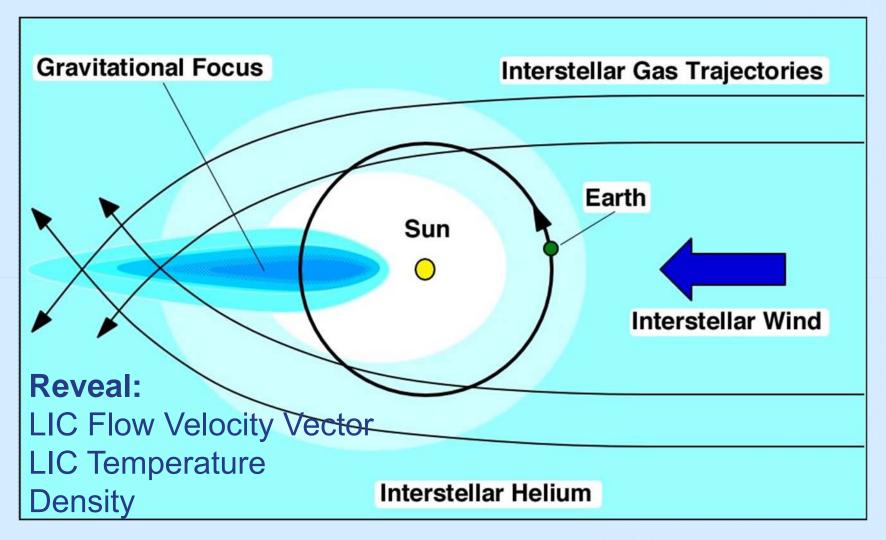


- Introduction
 Pickup He: Diagnostics of the Local Interstellar Medium
- Motivation
 Learn more about the causes of the large variability of Pickup lons
 Use pickup ions to study M/Q dependent injection and acceleration
 (CIRs, IP shocks)
- STEREO PLASTIC and Data Acquisition
 Data Analysis and Pickup Ion Spectra
- Three Cone Traversals in 2007 (A, B) and 2008 (A)
 Pickup Ion Fluxes and Solar Wind Parameters
- Summary & Conclusions





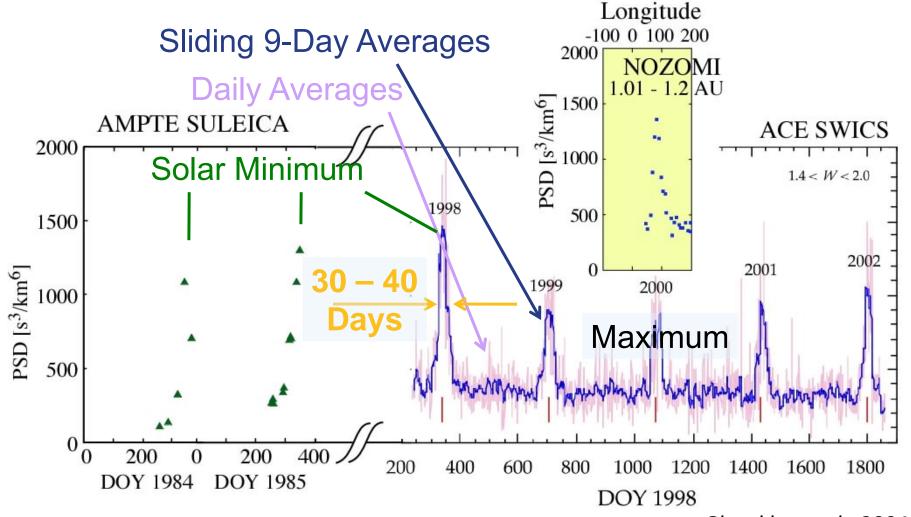
Interstellar He Cone Observations





Combined He Cone PUI Observations AMPTE, ACE & NOZOMI





Möbius et al., 1985

Gloeckler et al., 2004



Strong Variations in Pickup Ion Distributions



- Pickup Ion Distributions are Highly Variable
- Pickup Ions React to:
 - IMF Orientation (early Rollover)
 - IMF Strength (Compression and Rarefaction)
 - SW Density (Compression and Rarefaction)
 - Ionization rates and probably several unknown causes
- For Deduction of LIC Parameters Mitigation Sought with:
 - Long Integration Times

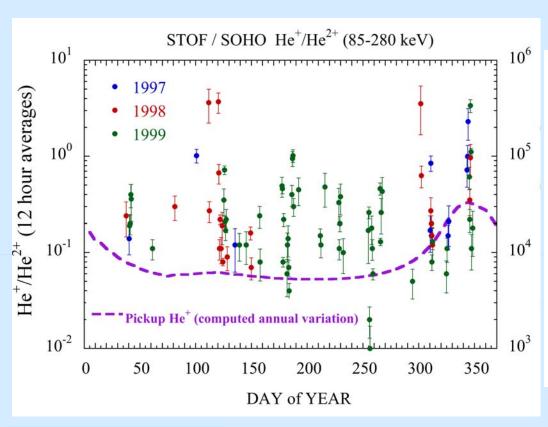
Better: Attempt to Understand Causes

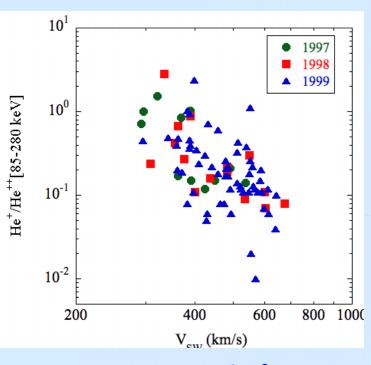
- Study of Ion Transport, Injection and Acceleration Processes (M/Q-dependent)
- STEREO (also ACE) Provide Simultaneous Observations



HE+ PICKUP IONS AT SUPRATHERMAL ENERGIES Events Related to Interplanetary Shocks







Anti-correlation of He⁺/He²⁺ with V_{SW}

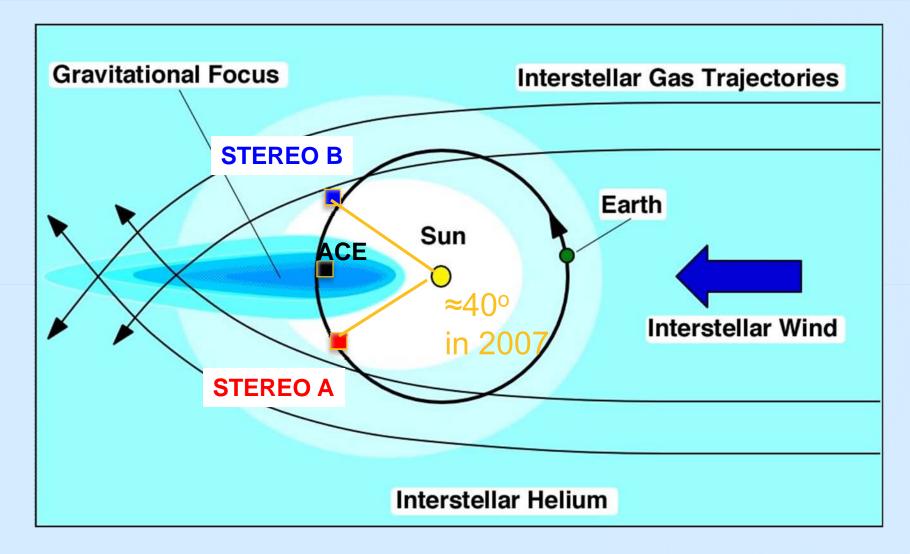
Hovestadt et al., 1984 Gloeckler, et al., 1994 Klecker et al., 2001 Kucharek et al., 2003

STEREO: High Sensitivity will enable study of individual events and time variations on short time scales (~ few hours)



Simultaneous Interstellar He Cone Observations

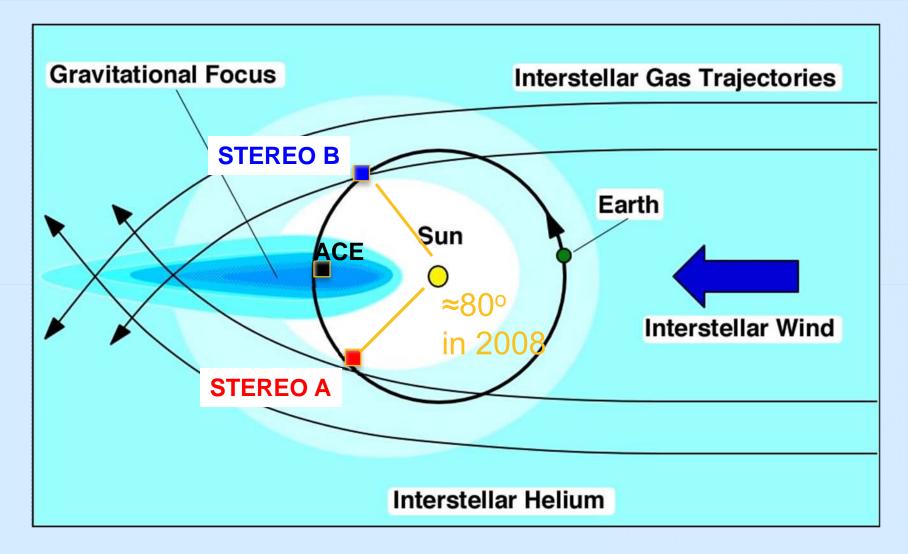






Simultaneous Interstellar He Cone Observations







He+ Analysis with STEREO PLASTIC

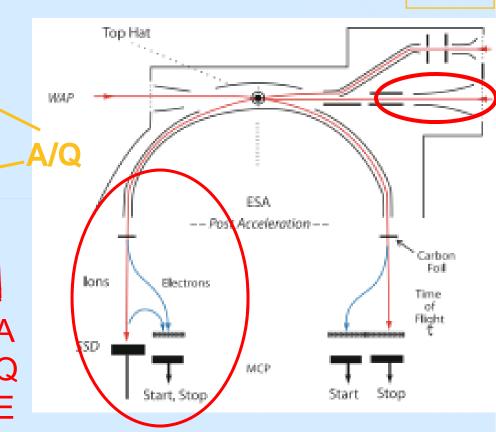
-> E/A



• STEREO PLASTIC Combines:

SUN

- Electrostatic Analyzer
- + Post-Acceleration -> E/Q
- Time-of-Flight
- Solid State Detectors



Data: Main Channel, Solar Wind Direction

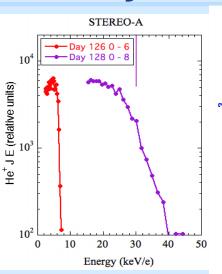
(Double + Triple Coincidences)

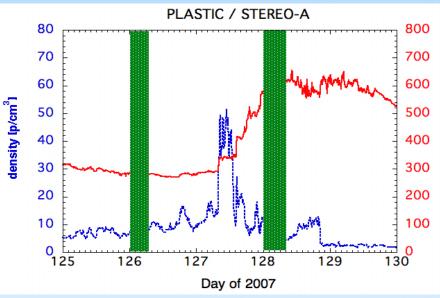


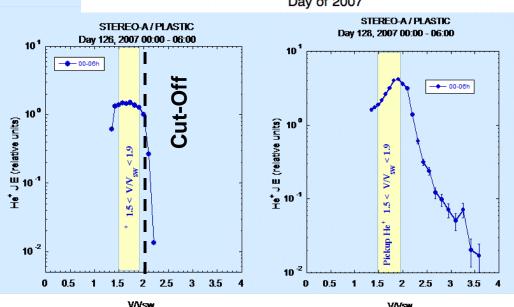


He+ Analysis: PUI Spectra

- Accumulate all **TOF Events with** and without SSD
- Counts Normalized with **Priority Rates**
- Sort in V/V_{sw} **Natural Frame for PUI Spectra**
- Average over 24 h and $1.5 < V/V_{sw} < 1.9$







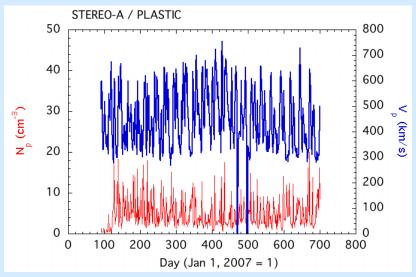
V/Vsw

V/Vsw



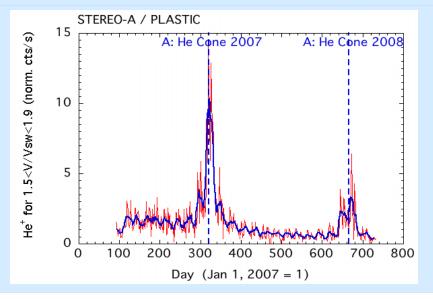
STEREO-A OBSERVATIONS IN 2007/2008





Solar Wind in 2007 and 2008:

High Speed Streams and CIRs



Pickup He+

- Large variability of daily averages
- 9-day Sliding Average Smoothes some of the known strong variations

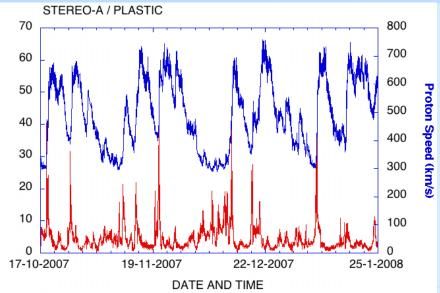


Proton Density (p/cm³)

He⁺ for 1.5<V/Vsw<1.9 (norm. cts/s)

STEREO-A OBSERVATION OF FOCUSING CONE IN 2007



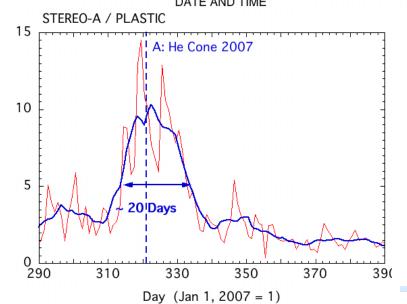


Solar Wind in 2007 and 2008:

High Speed Streams and CIRs

Pickup He+

- Focusing Cone Fine Structure
 Influenced by High- Speed Streams
- Focusing Cone FWHM ~20 days

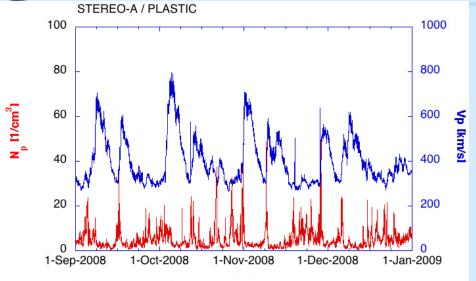




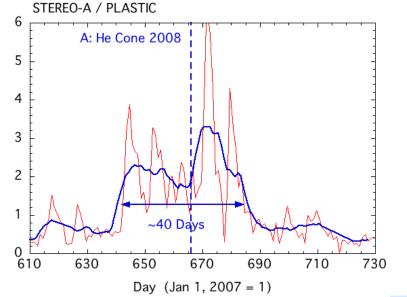
He⁺ for 1.5<V/Vsw<1.9 (norm. cts/s)

STEREO-A OBSERVATION OF FOCUSING CONE IN 2008





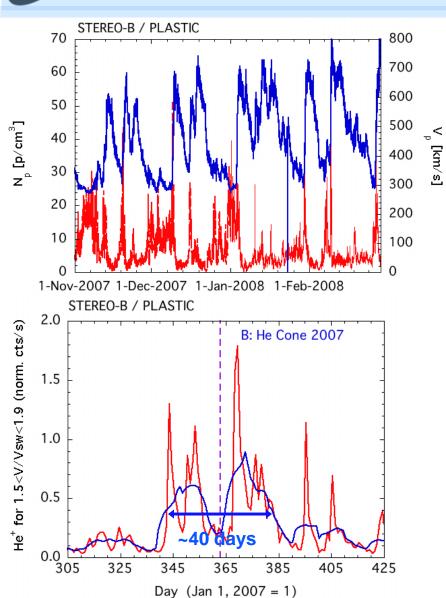
Solar Wind in 2008



Pickup He+

 Focusing Cone Fine Structure and Large Scale Structure Strongly influenced by High-Speed Streams

STEREO-B OBSERVATION OF FOCUSING CONE IN 2007/2008



Solar Wind in 2007 and 2008:

High Speed Streams and CIRs

Pickup He+

- Focusing Cone Large Scale
 Structure Strongly Influenced by
 High-Speed Streams
- Focusing Cone FWHM ~40 days



Summary and Conclusions



- Three Interstellar He Cone Traversals with STEREO A & B in 2007 and 2008
- Strong Daily Variations of Pickup Ion Distributions
- After Smoothing with 9-Day Sliding Average the 3 Cone Traversals Appear Substantially Different
- A 2007: Focusing Cone Fine Structure Influenced by High- Speed Streams
- B 2007 and A 2008: Large Scale Structure of Focusing Cone Influenced by Solar Wind Streams

Next to do:

- Convert normalized counts into flux using (time variable) efficiencies
- Correlation of He $^+$ variability on A and B, correlation with V_p , N_p , B, etc
- Investigate Suprathermal tails of H⁺, He²⁺, and He⁺ in CIRs
 - and in the CME-related IP shocks we all are waiting for