

Heliospheric Imager – Scientific Operations

- HI Operations Document – R. Harrison
- HI Image Simulation – C. Davis & R. Harrison
- HI Operations Scenarios – R. Harrison & S. Matthews
- HI Beacon Mode Specification – S. Matthews

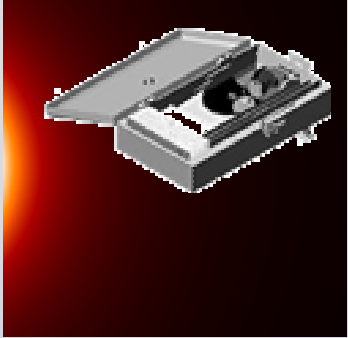


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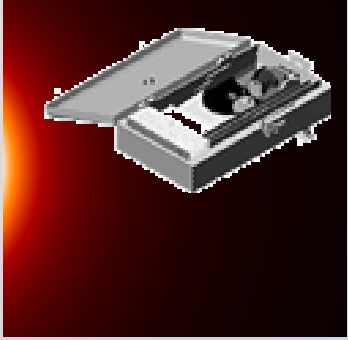
Heliospheric Imager – Scientific Operations

➤ HI in a nutshell:

First opportunity to observe Earth-directed CMEs along the Sun-Earth line in interplanetary space - the first instrument to detect CMEs in a field of view including the Earth!

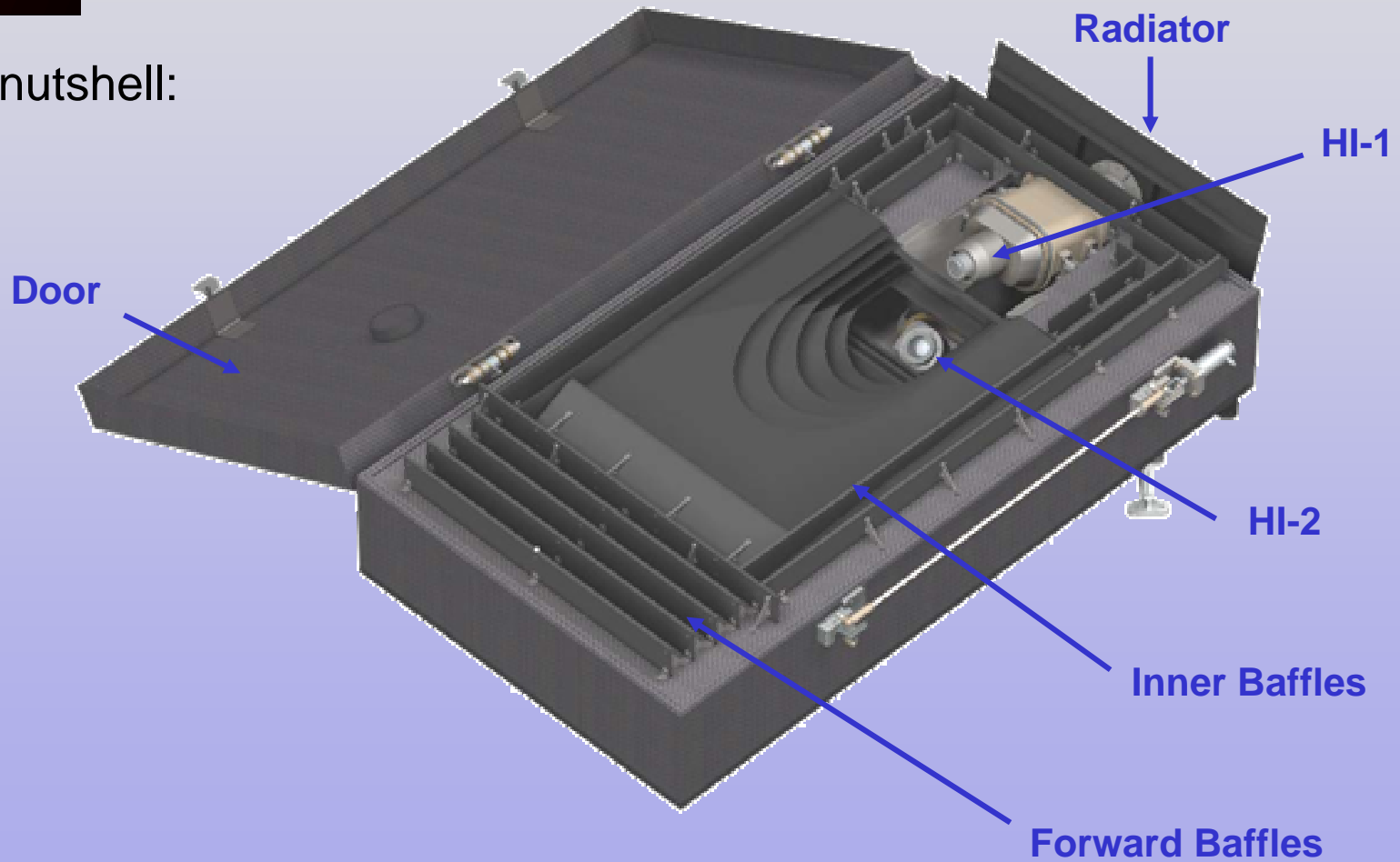
First opportunity to obtain stereographic views of CMEs in interplanetary space - to investigate CME structure, evolution and propagation.

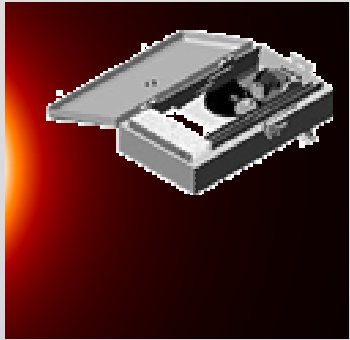
Method: Occultation and baffle system, with wide angle view of the heliosphere, achieving light rejection levels of 3×10^{-13} and 10^{-14} of the solar brightness.



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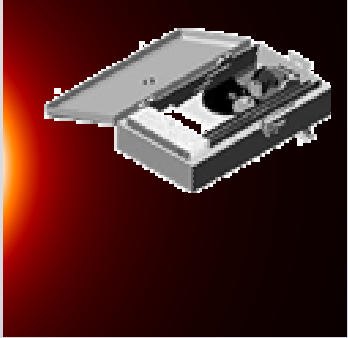




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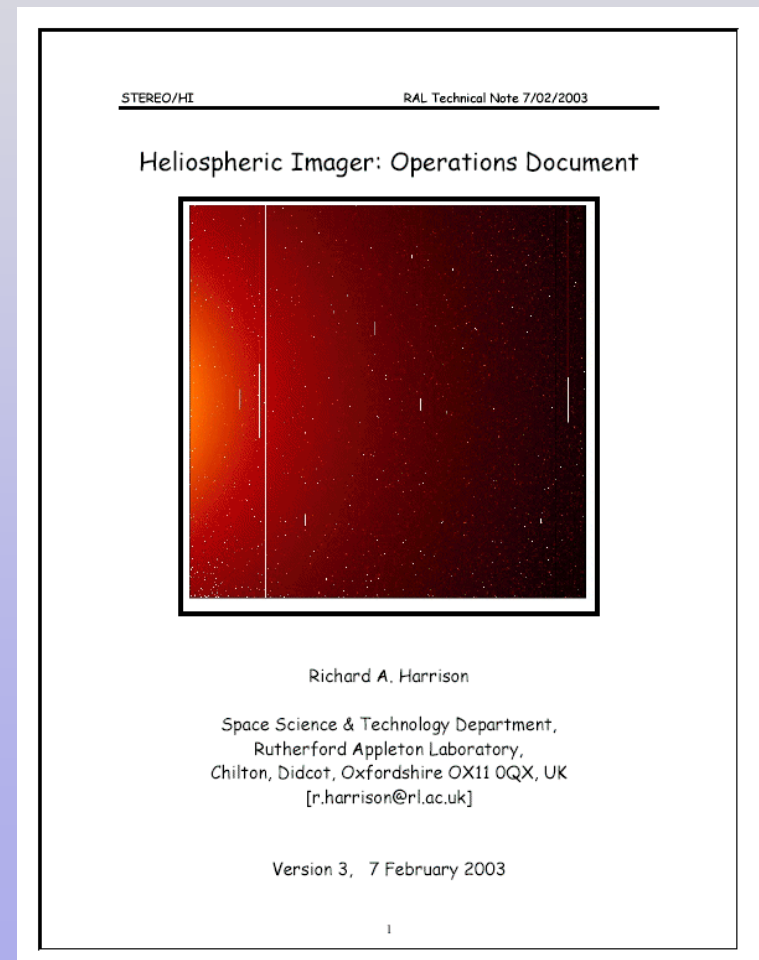
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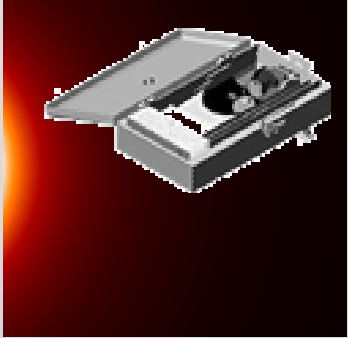
	HI-1	HI-2
Instrument Type	Externally-Occulted Coronagraph	Externally-Occulted Coronagraph
Centre of Field-of-View Direction	Along Sun-Earth Line $\theta = 13.65 \text{ deg}$	Along Sun-Earth Line $\theta = 53.35 \text{ deg}$
Angular Field-of-View	20 deg	70 deg
Coronal Coverage	12 - 84 R_{sun}	66 - 318 R_{sun}
Overlap With COR2	12 - 15 R_{sun}	N/A
Overlap With HI-1	N/A	66 - 84 R_{sun}
Baseline Image (2 x 2 Binning)	1024 x 1024	1024 x 1024
Image Pixel Scale (Binned)	70 arcsec	4 arcmin
Spectral Bandpass	630 - 730 nm	400 - 1000 nm
Exposure Time	12 - 20 sec	60 - 90 sec
Nominal Images Per Sequence	70	50
Required Cadence (Per Sequence)	60 min	120 min
Brightness Sensitivity	$3 \times 10^{-15} B_{\text{sun}}$	$3 \times 10^{-16} B_{\text{sun}}$
Straylight Rejection	$3 \times 10^{-13} B_{\text{sun}}$	$10^{-14} B_{\text{sun}}$
Brightness Accuracy	10%	10%



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- HI Operations Document
 - HI Operations Document – Version 4 released Dec 1, 2003
 - Author: Richard Harrison, HI Principal Investigator
 - Document located at UK Web site: <http://www.stereo.rl.ac.uk>
 - The HI team is not aware of any other instrument operations document on STEREO.





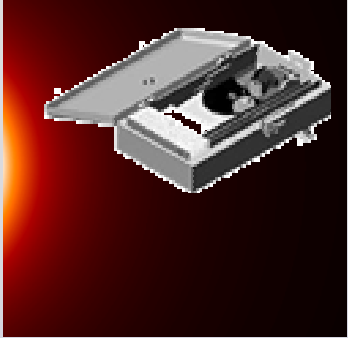
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➤ HI Operations Document

➤ **Purpose:** Sets out plans for the operation of the Heliospheric Imager. It is intended that this information be used as an input to the discussion on

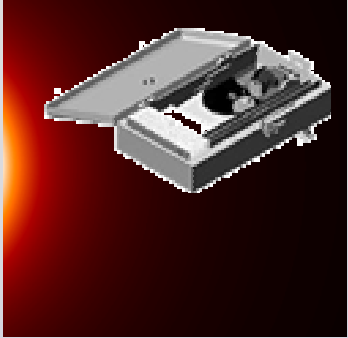
- the development of on-board and ground software (including planning tool software, archive software and data handling, inspection and analysis software),
- payload operations planning,
- commanding,
- monitoring and data receipt,
- data handling and archiving.

In short – it spells out the requirements on operation and software.



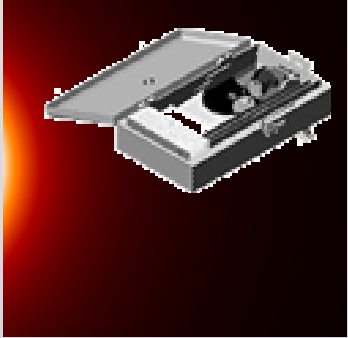
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- HI Operations Document – contents:
 - Operations planning and implementation
 - HI Scientific operation
 - Data monitoring and archiving
 - Image processing and calibration requirements
 - Instrument monitoring and maintenance
 - Commissioning plan
 - The beacon mode
 - Software requirements
 - Scientific operations sequences



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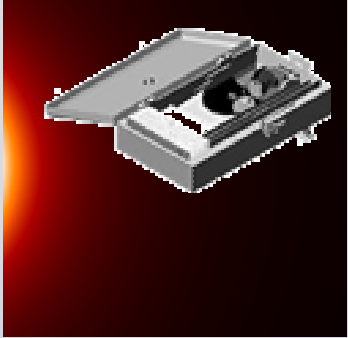
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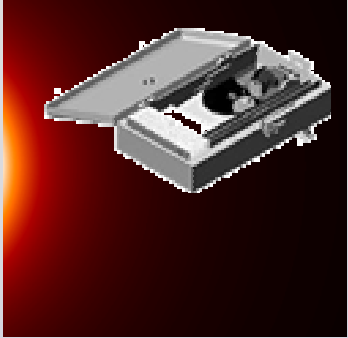
- With regard to software and operations requirements, the HI Operations Document lists 34 requirements which must be considered by the SECCHI software team and those planning the operations facilities.
- These requirements range from flexibility of programming parameters such as exposure times, to the return of partial frames, from cosmic ray cleaning to the definition of the beacon mode.



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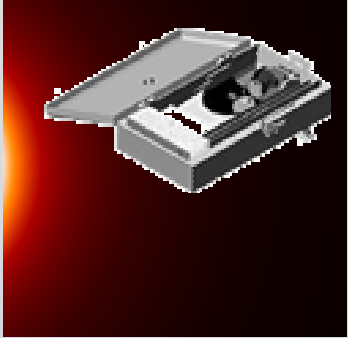
➤ HI Operations Scenarios:

- With regard to HI Scientific Operations Sequences, we have continued the design of specific operations schemes, aimed at addressing specific scientific questions.
- This is used to define the operation and its flexibility and comes out of the highly successful ‘Blue Book’ studies of CDS/SOHO.
- The products are a clear understanding of how we wish to use the instrument, and clear definitions of the requirements on software and operations.
- 15 scenarios so far – next slide...



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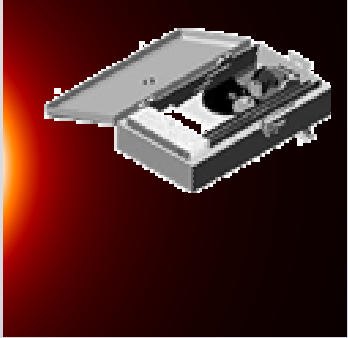
<i>Study</i>	<i>Author</i>
Synoptic CME programme	R. Harrison
Beacon mode	Matthews, Harrison, Davis
Impact of CME on Earth	R. Harrison
Understanding how observations at L1 & SECCHI are related	P. Cargill
CMEs in interplanetary space	P. Cargill
3-D structure of interplanetary CMEs	L. Green
CME onset	S. Matthews
Particle acceleration at CME shocks	S. Matthews
The relationship between CMEs and magnetic clouds	S. Matthews
Boundary regions between fast & slow streams in the solar wind	A. Breen
Development of co-rotating interaction regions	A. Breen
Solar wind microstructure	A. Breen
Differential drift velocities in the fast & slow solar winds	A. Breen
Remote solar wind measurements from 3-D obs. of cometary ion tails	G. Jones
Interplanetary acceleration of ICMEs	M. Owens



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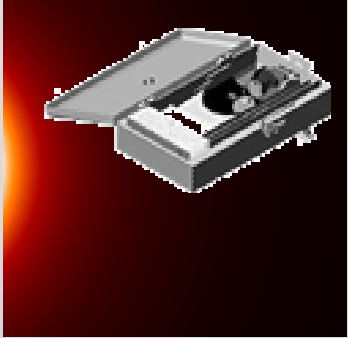
➤ HI Scientific Operations Scenarios – the Synoptic Mode:

	HI -1	HI -2
Image array	1024x1024 (2kx2k summed)	1024x1024 (2kx2k summed)
FOV	20° (3.65-23.65)	70° (18.35-88.35)
Nominal Exposure	12 s	60 s
Summed Exposures	70	60
Synoptic Cadence	1 hr	2 hr
Telemetry Rate	2.9 kbit/s	1.5 kbit/s



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- The Beacon Mode –
 - Provided for quick data receipt, for space weather purposes
 - HI is a key player in this – the only instrument to see CMEs with Earth within the boundary of the FOV
 - Options:
 - Reduced resolution images;
 - N-S strip Sunward of Earth;
 - Partial images.



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➤ The Beacon Mode –

➤ Current plan:

Returned image	128x128 pixel image (summed from 2048x2048 array on board)
Rate	1 image per hour, alternately HI-1 and HI-2.
Pixel depth	32 bits (defined by on board summed data)
Nominal telemetry	147 bit/sec.

Note: The beacon mode must be programmable so we can explore different approaches particularly in the early mission.